

Super-Massive Black Holes in Compact Galaxies

Remco van den Bosch

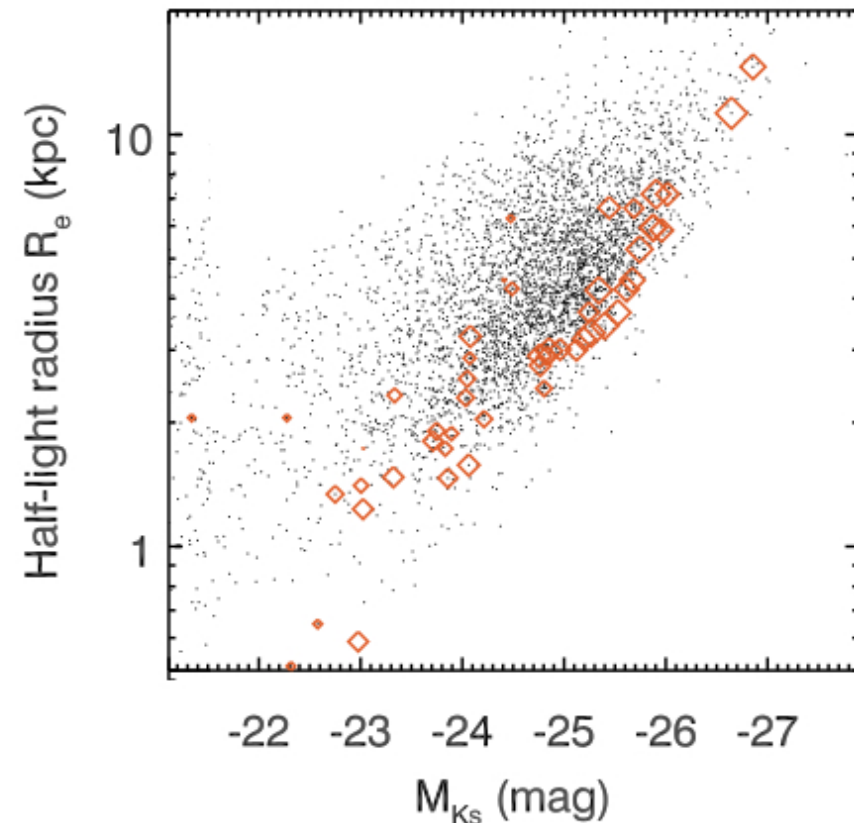
MPIA

Aaron Barth
Karl Gebhardt
Kayhan Gültekin
Bernd Husemann
Ronald Läsker
Mariya Lyubenova
Akin Yildirim
Glenn van de Ven
Jonelle Walsh



DIRECT BLACK HOLE MASSES

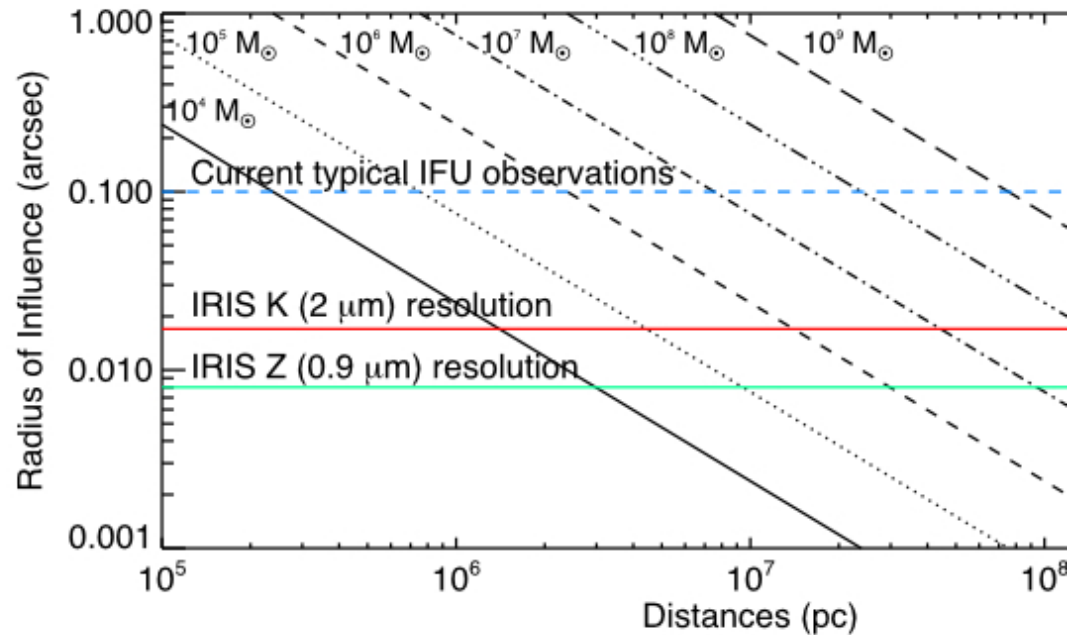
- The direct black hole masses in nearby galaxies are the basis for all other BH mass estimates.
- Only ~80 have been measured to date.
- Requires high spatial resolution spectroscopy
ELT (Do+14), ALMA
Davis14)



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$$R_{\text{soi}} = \frac{GM_{\bullet}}{D\sigma^2} \propto \frac{\sigma^{2.2}}{D}$$





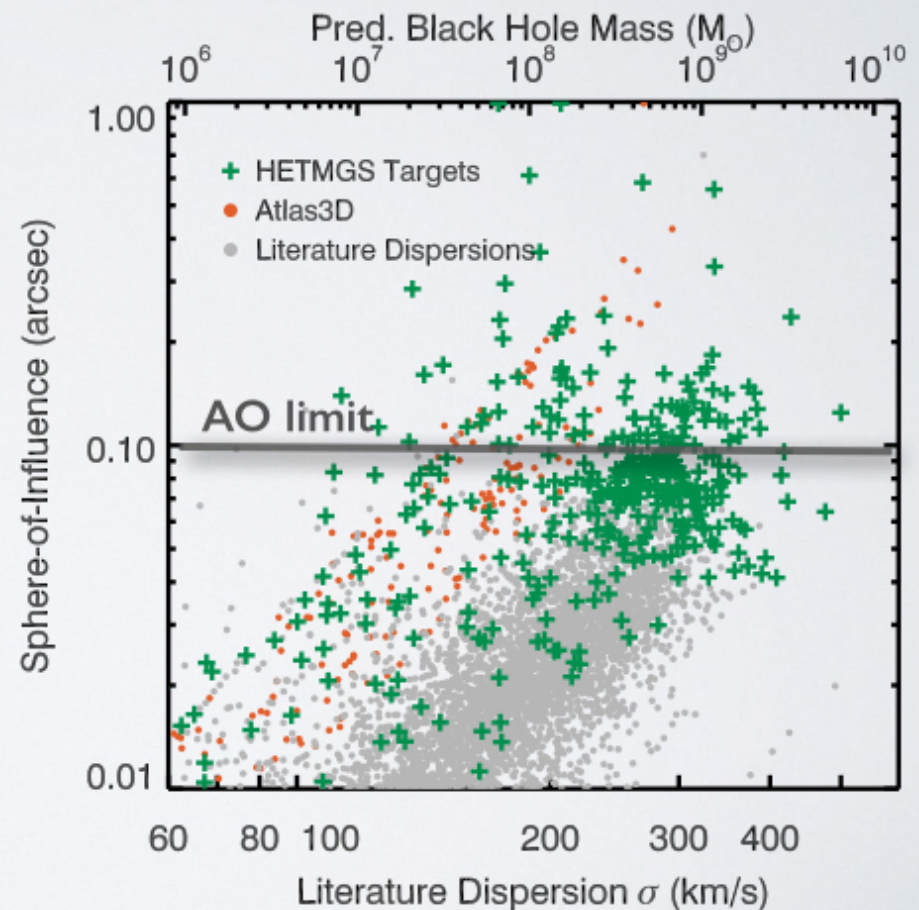
HET SURVEY

- Long slit spectra with the Marcario Low Resolution Spectrograph
- 4200-7400 Å, 106 km/s resolution, 1"×2.5' slit
- 1000 galaxies
- Distances less than ~ 140 Mpc
- Targeting the galaxies with the largest sphere of influences.
- Effectively probing the massive nearby galaxies



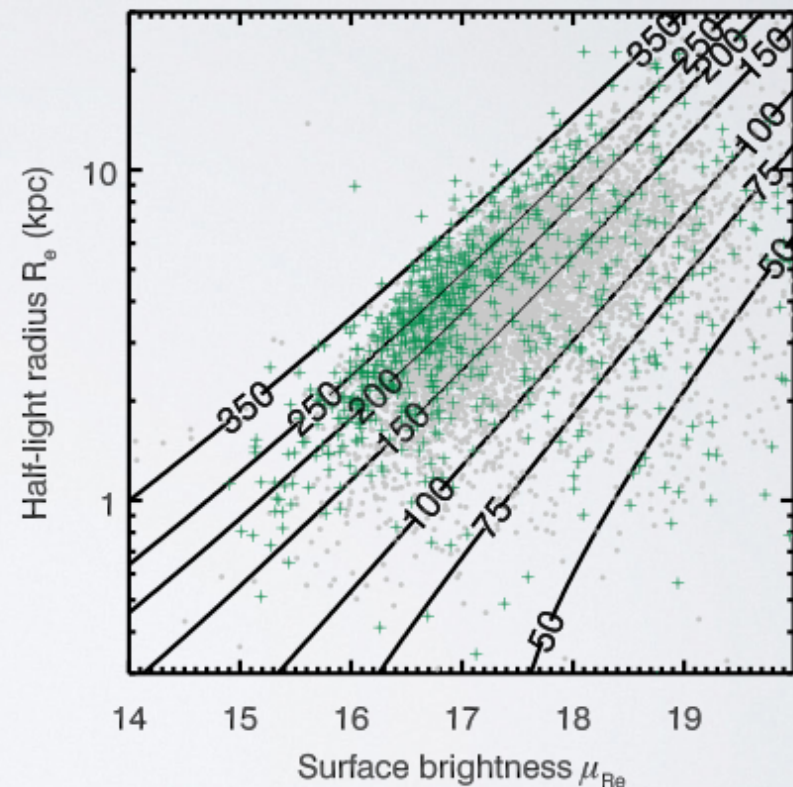
HET MASSIVE GALAXIES SURVEY

- Select candidate galaxies using literature velocity dispersion from Hyperlede database
- Predict black hole mass using M - σ
- Few targets with $\text{SOI} > 0.1''$
- Most nearby galaxies are not in SDSS



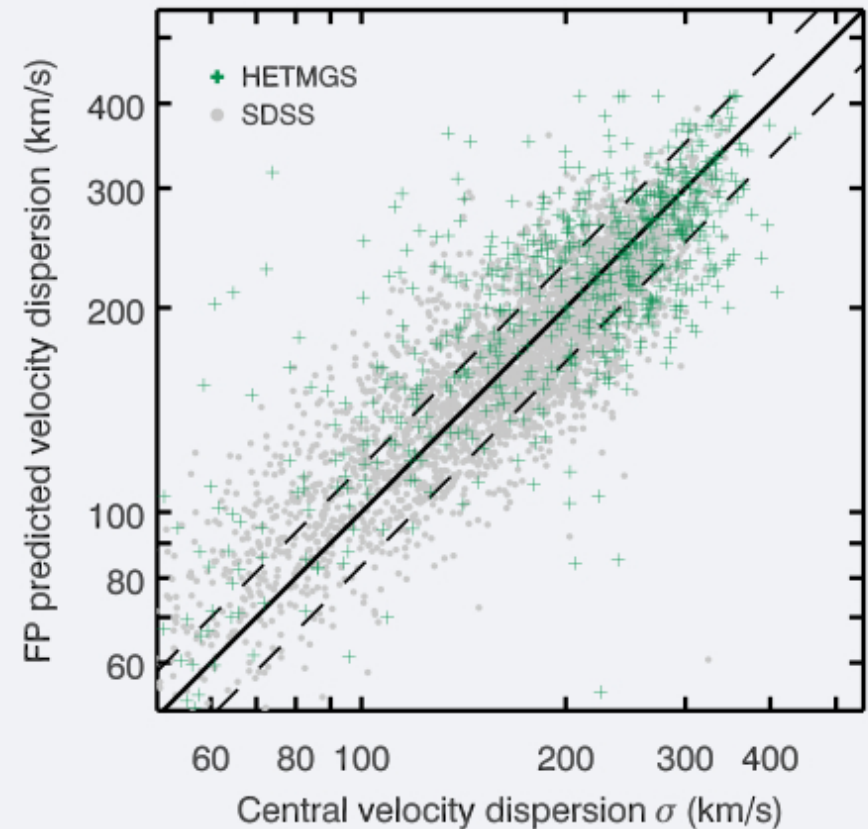
2MASS FUNDAMENTAL PLANE

- **Sample across the face of the FP**
- Not all nearby galaxies have dispersions
- Predict dispersion:
 - 2Mass XSC (Jarret+00)
 - 2MRS Redshifts (Huchra+12)
 - Fundamental Plane (Dressler+87)
 - No selection on galaxy type
- Dispersion estimates accurate up to 0.09 dex

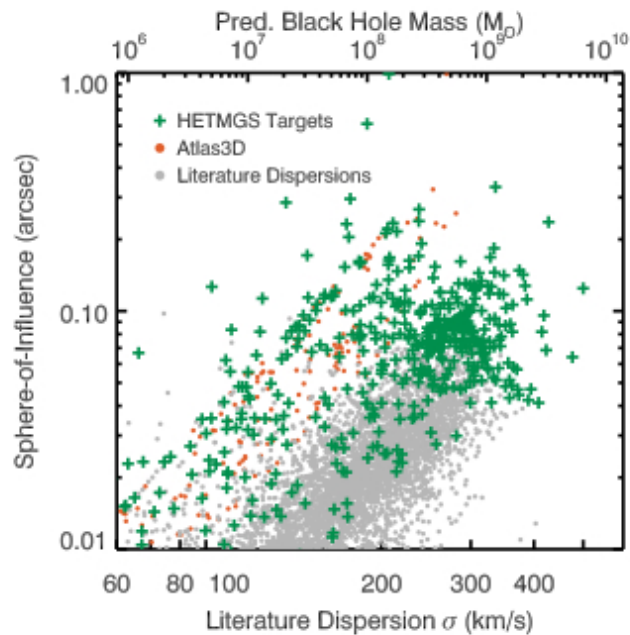


2MASS FUNDAMENTAL PLANE

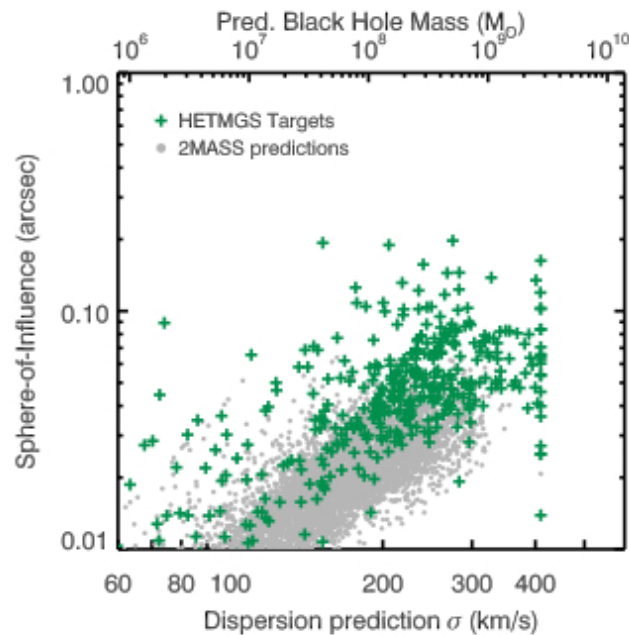
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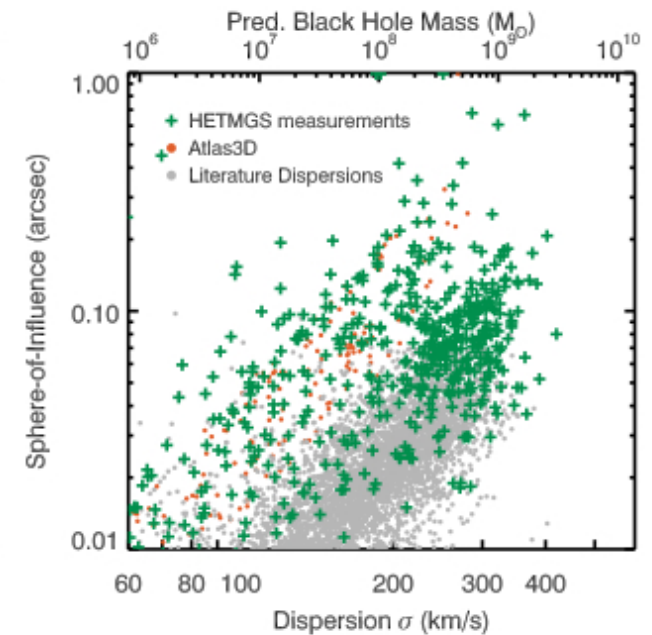
No literature dispersions



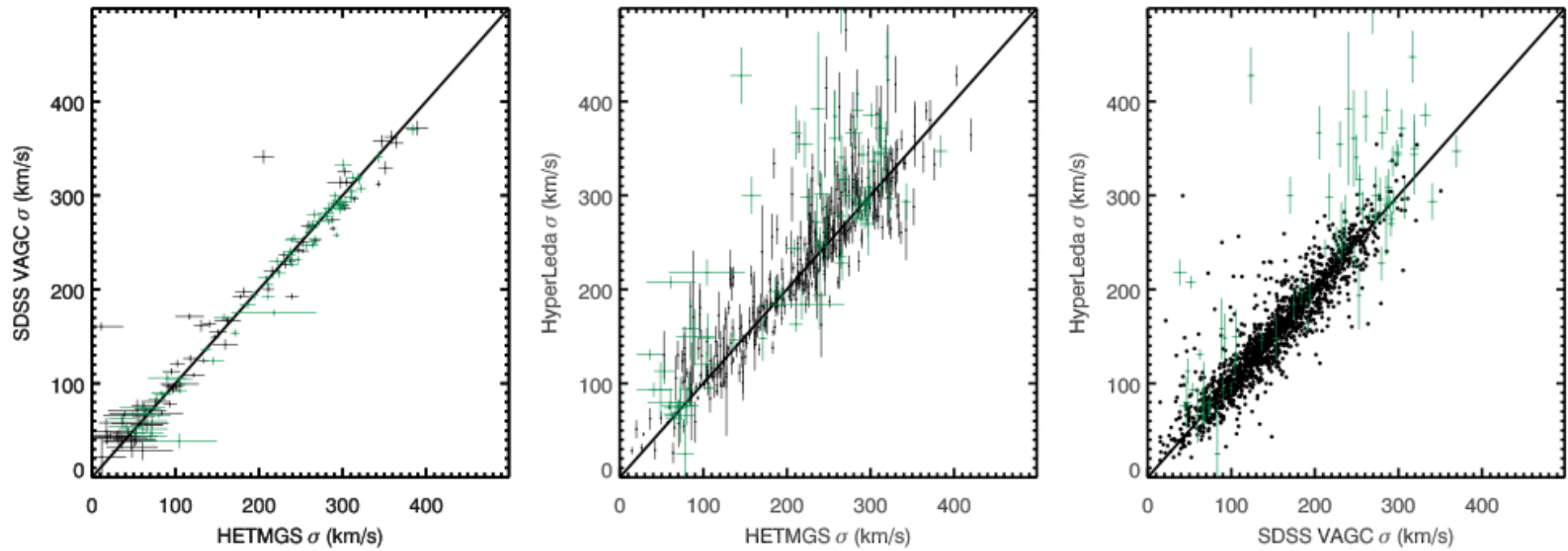
No literature dispersions



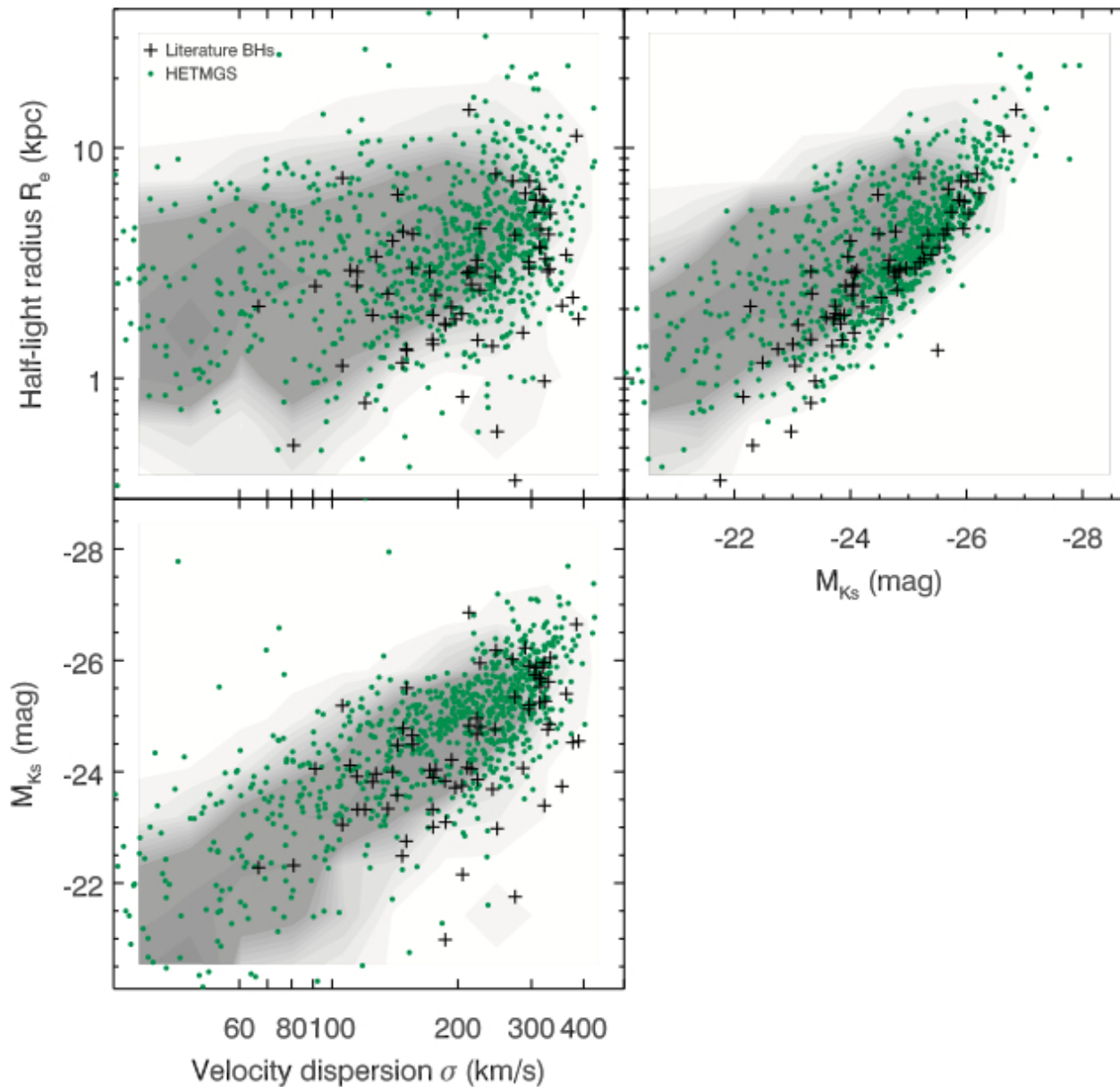
Final Survey



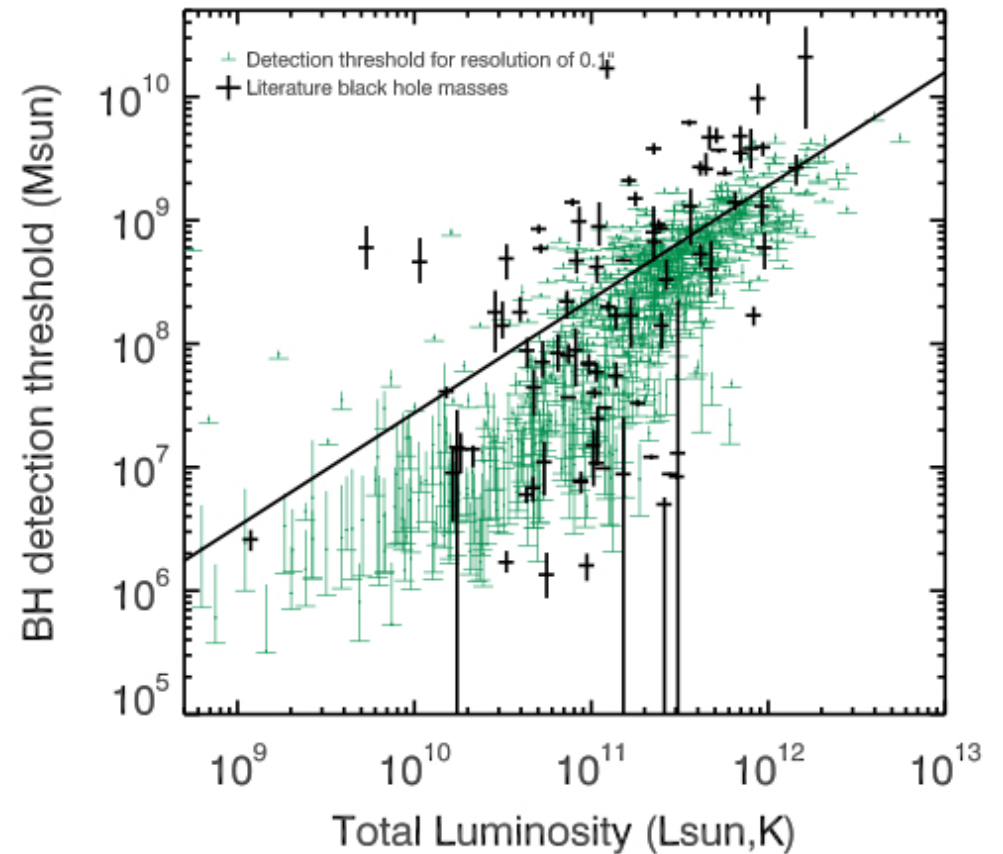
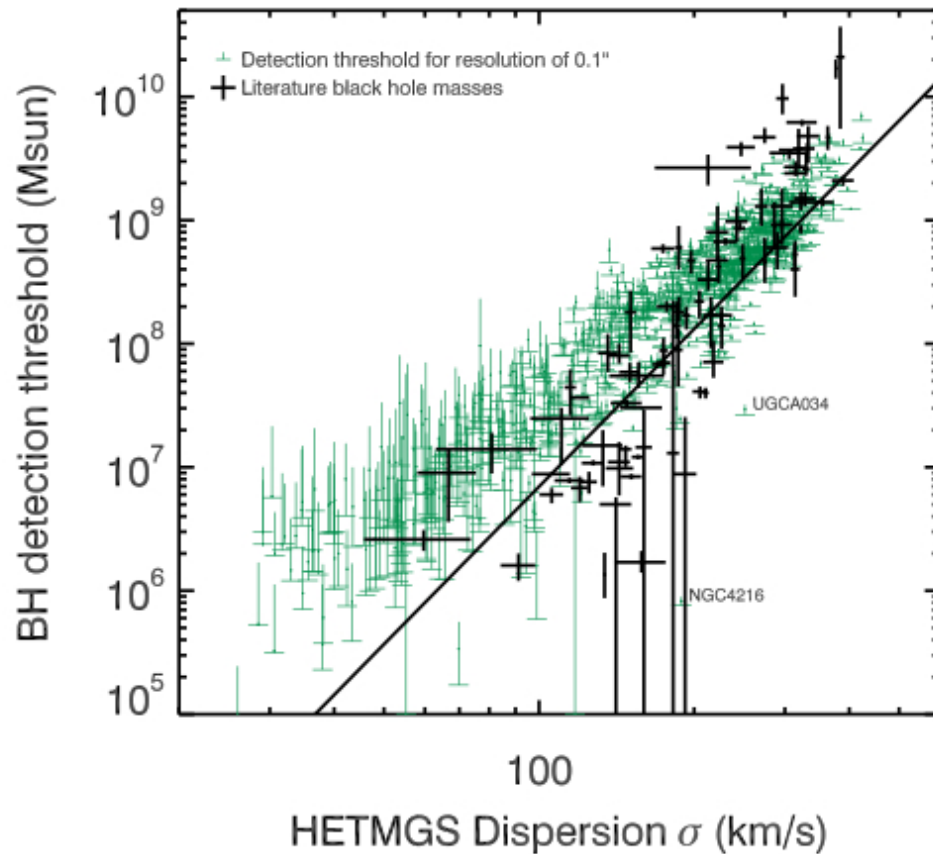
Survey completeness



Dispersion comparison



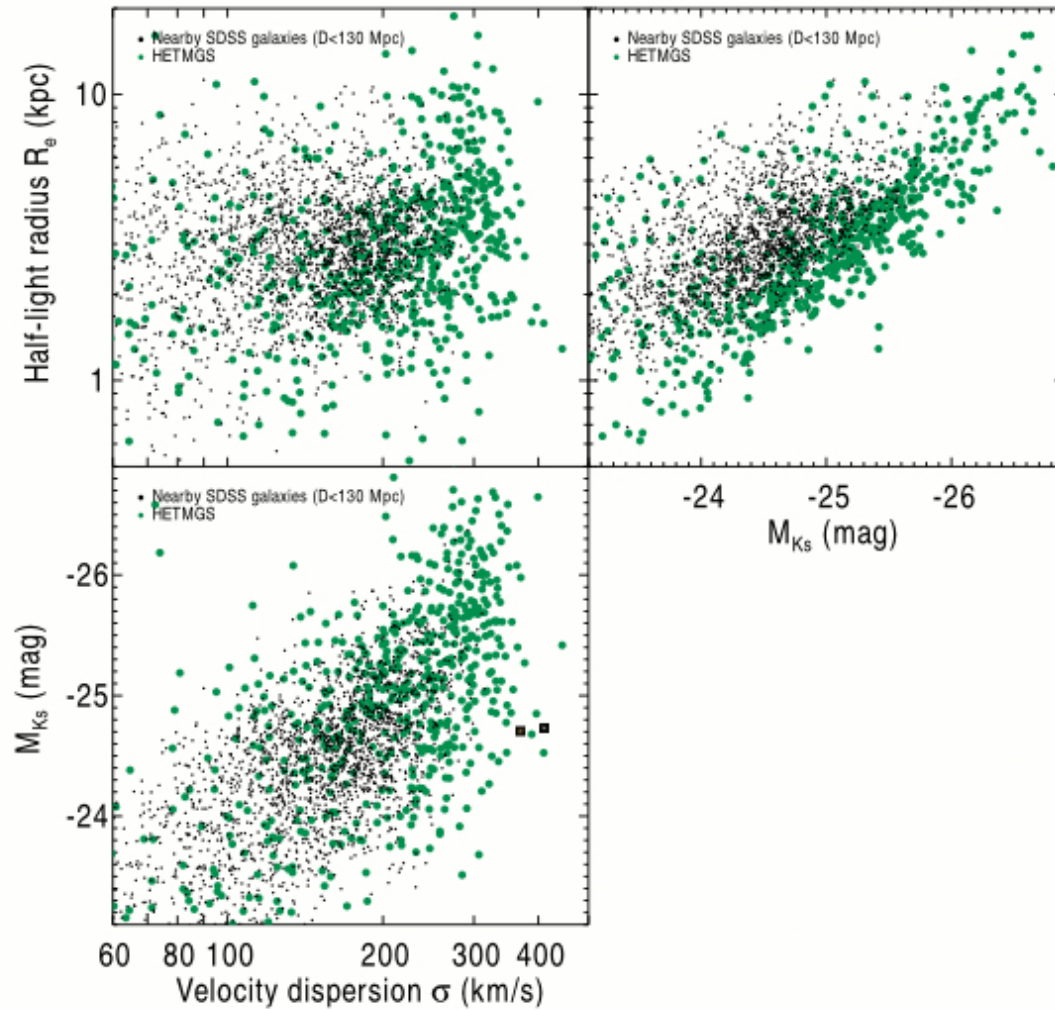
Distribution of BH hosts



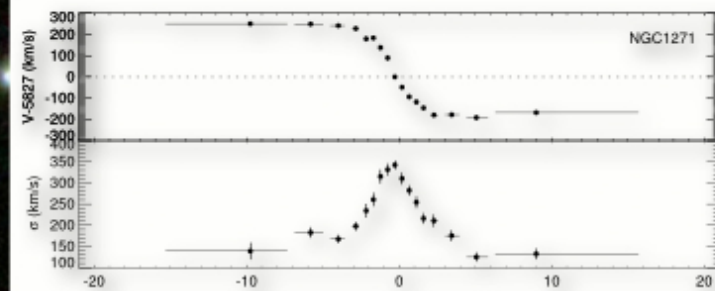
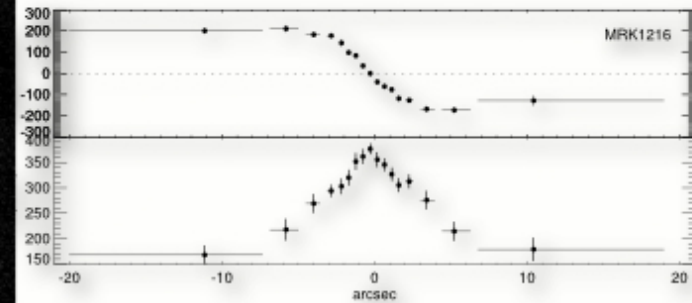
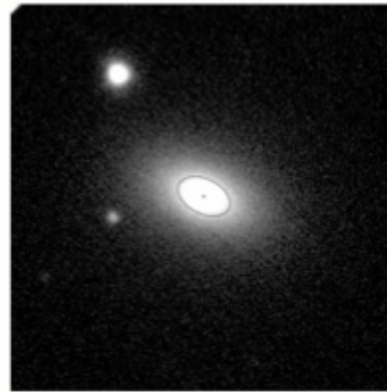
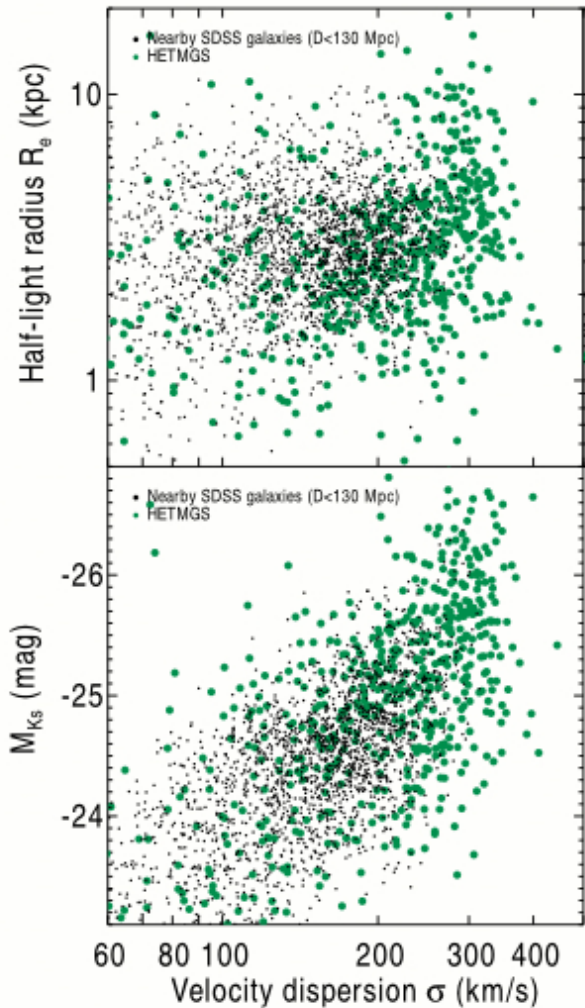
What can we expect?

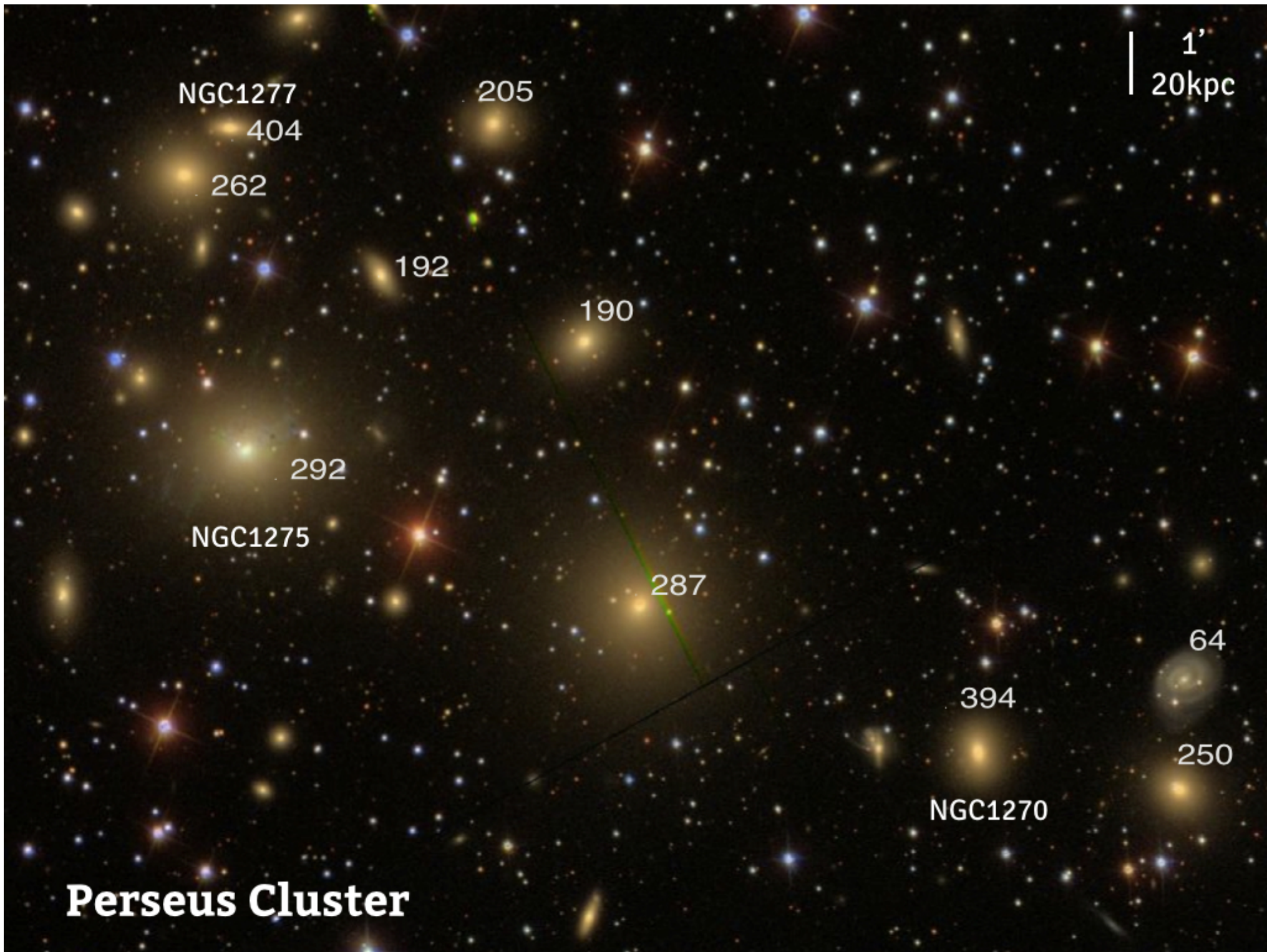
$$R_{soi} = \frac{GM_{\bullet}}{D\sigma^2}$$

COMPACT GALAXIES

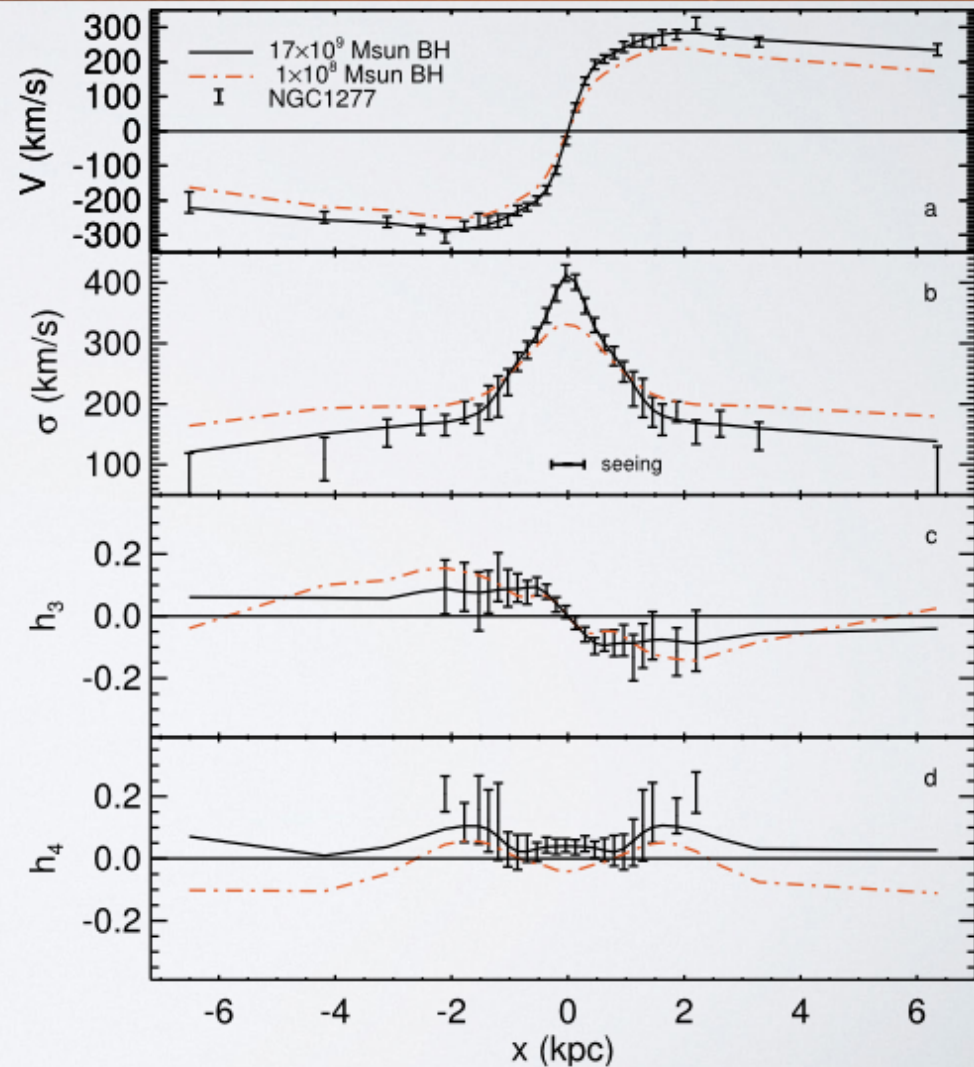
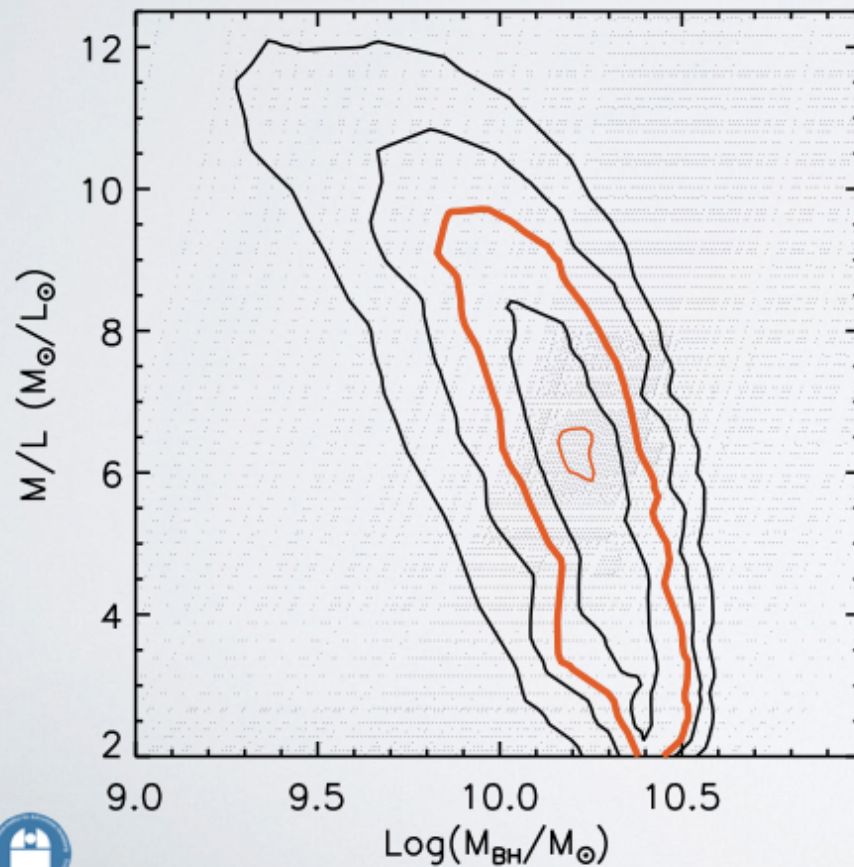
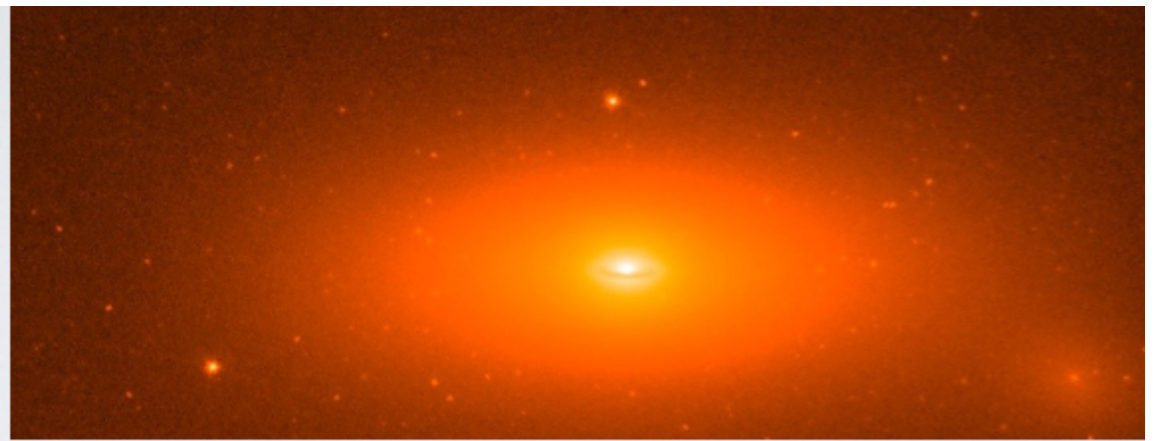


COMPACT GALAXIES

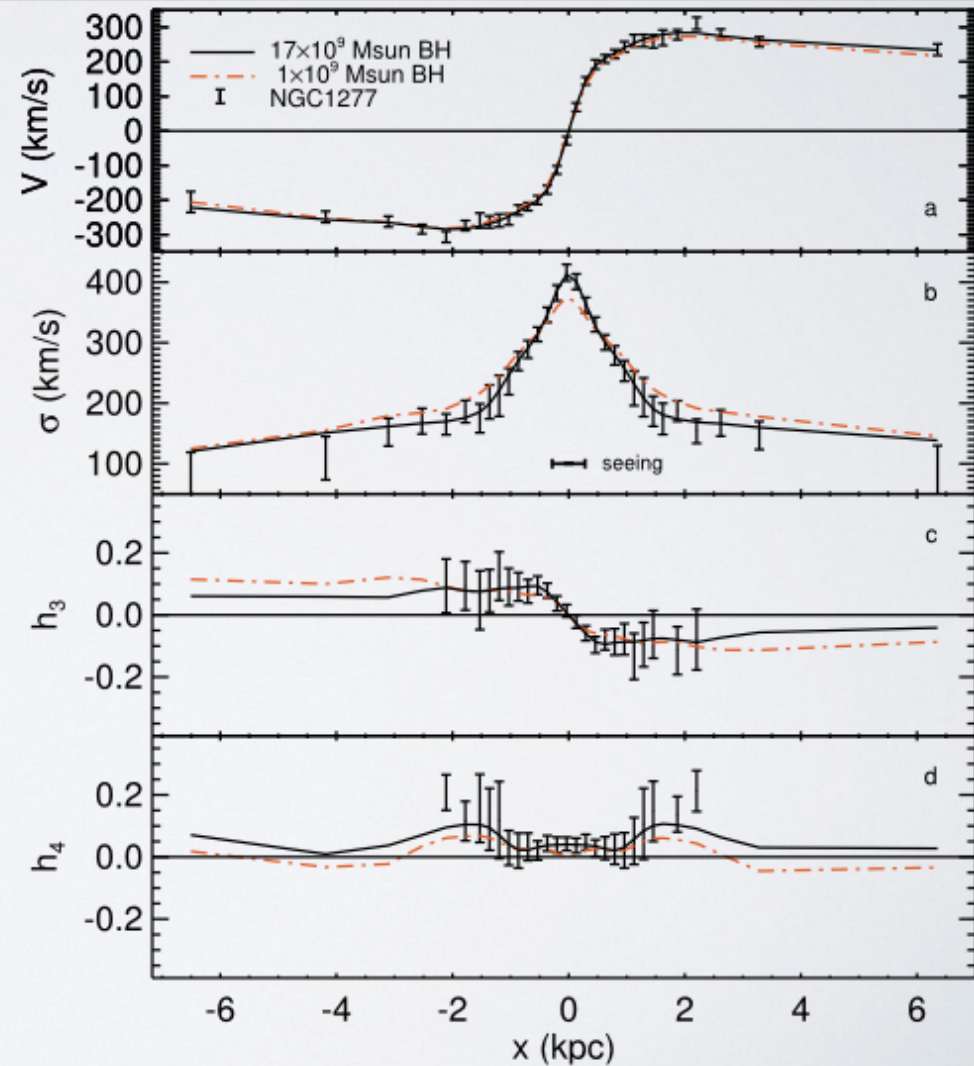
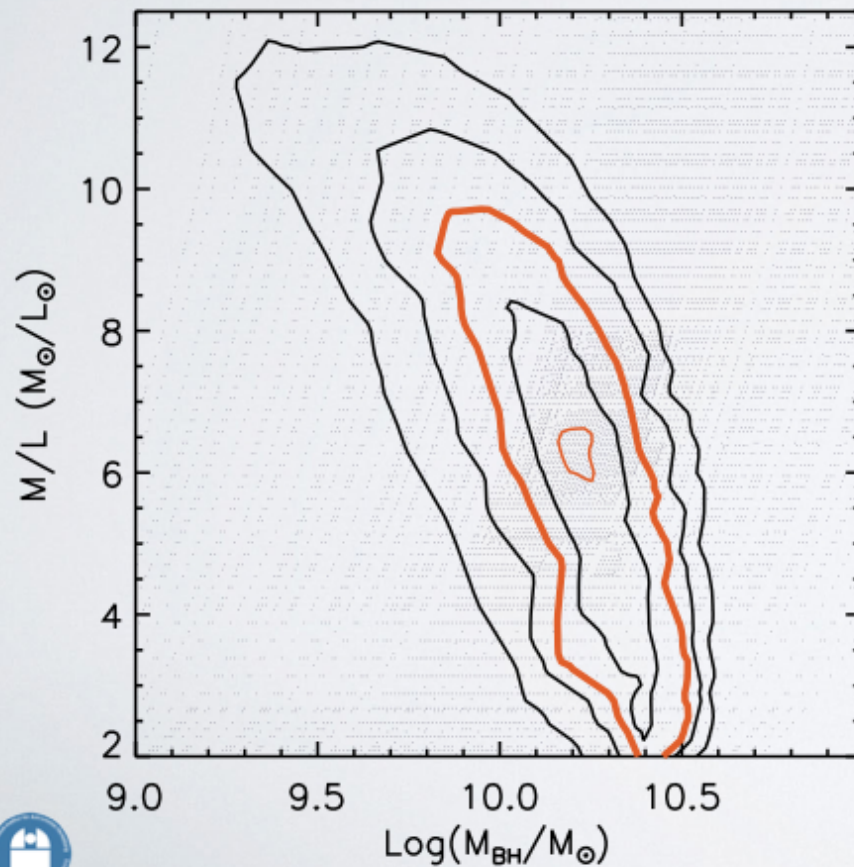
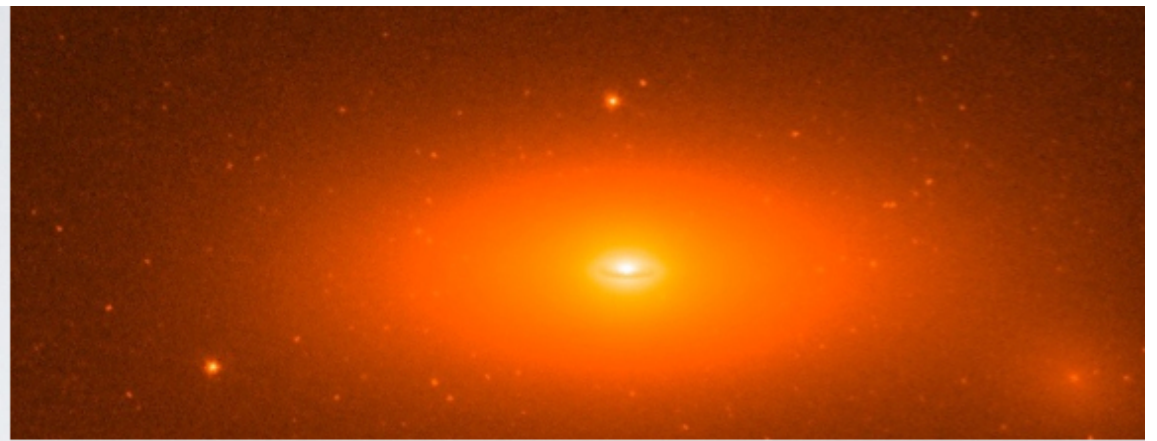




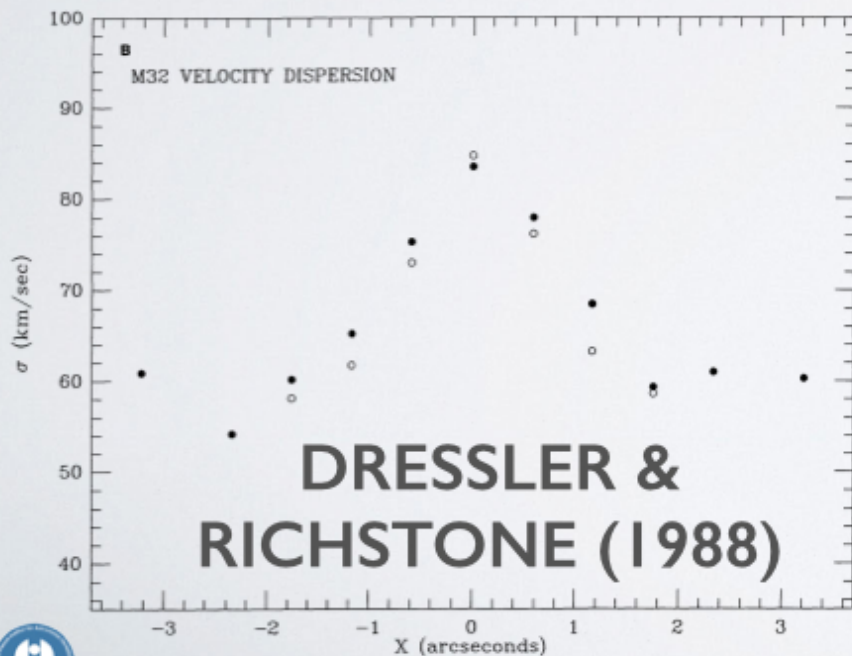
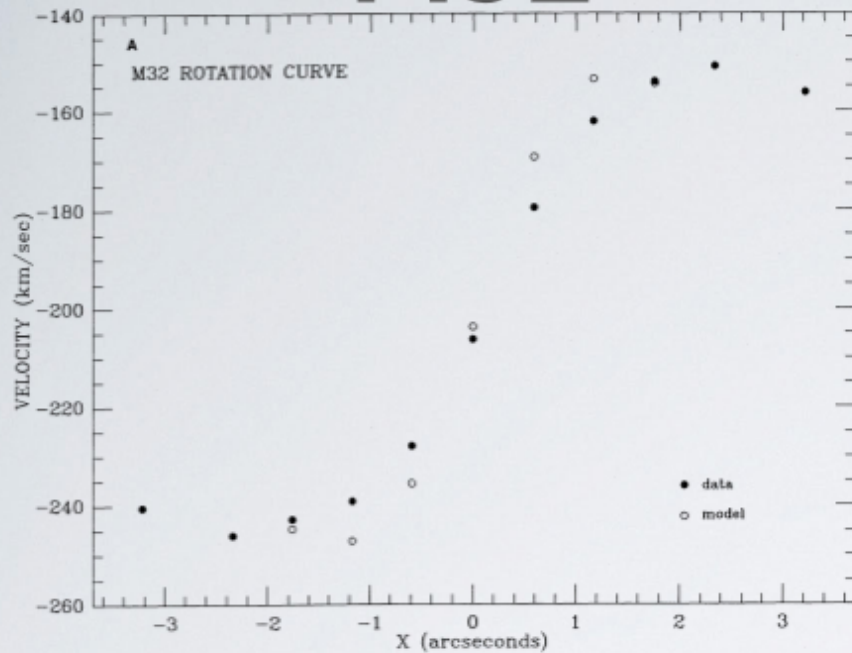
NGC 1277: A BIG BLACK HOLE IN A SMALL GALAXY



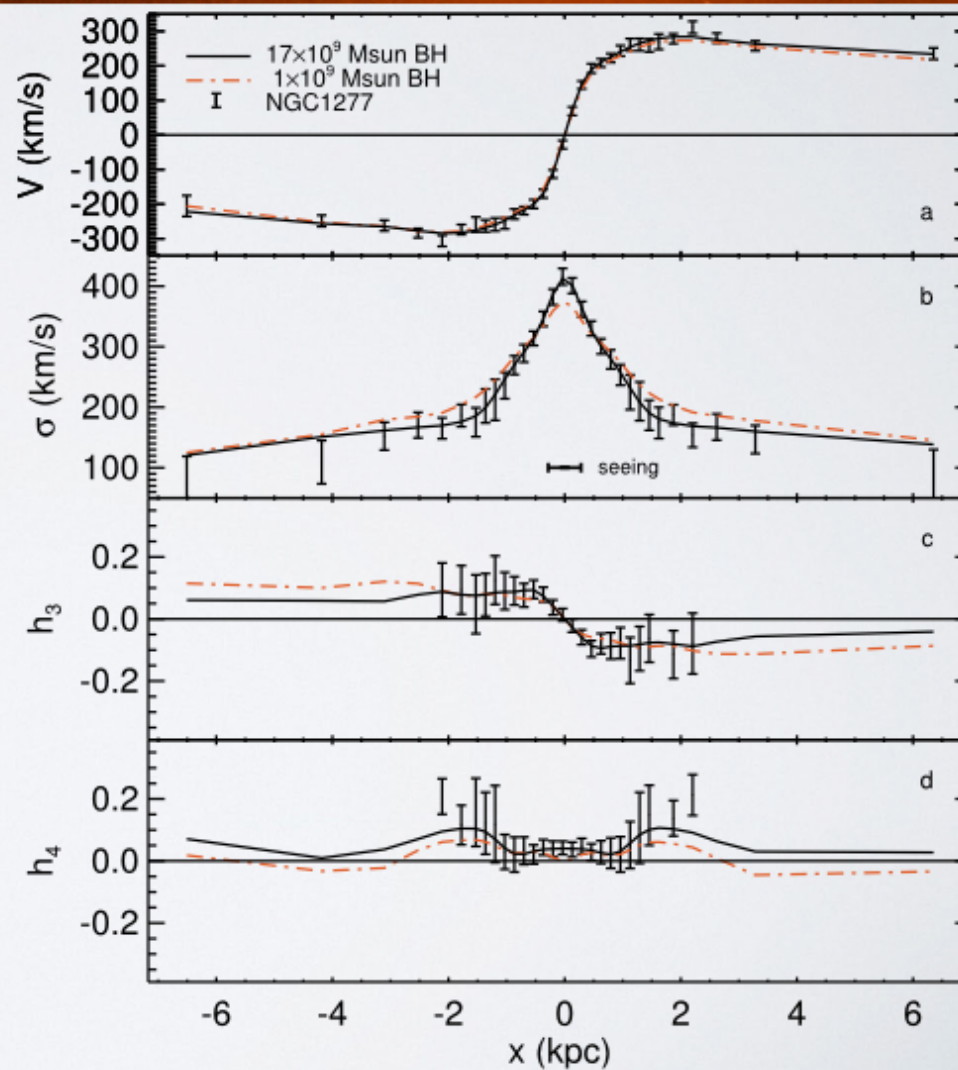
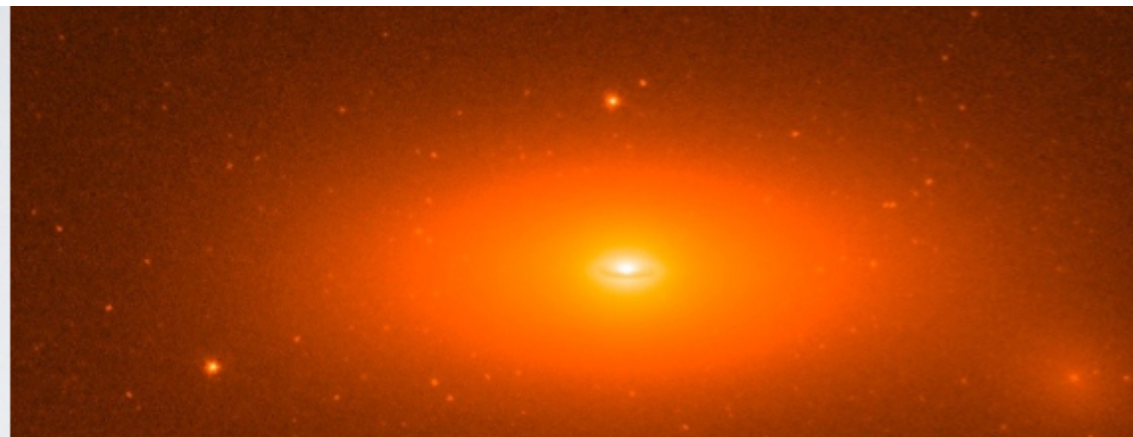
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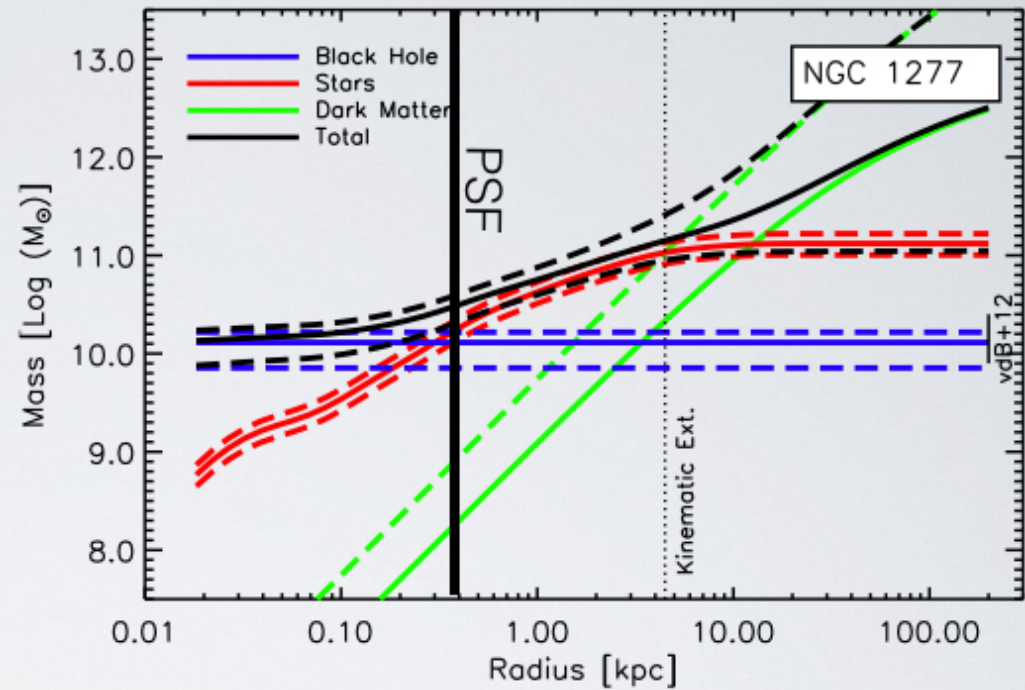
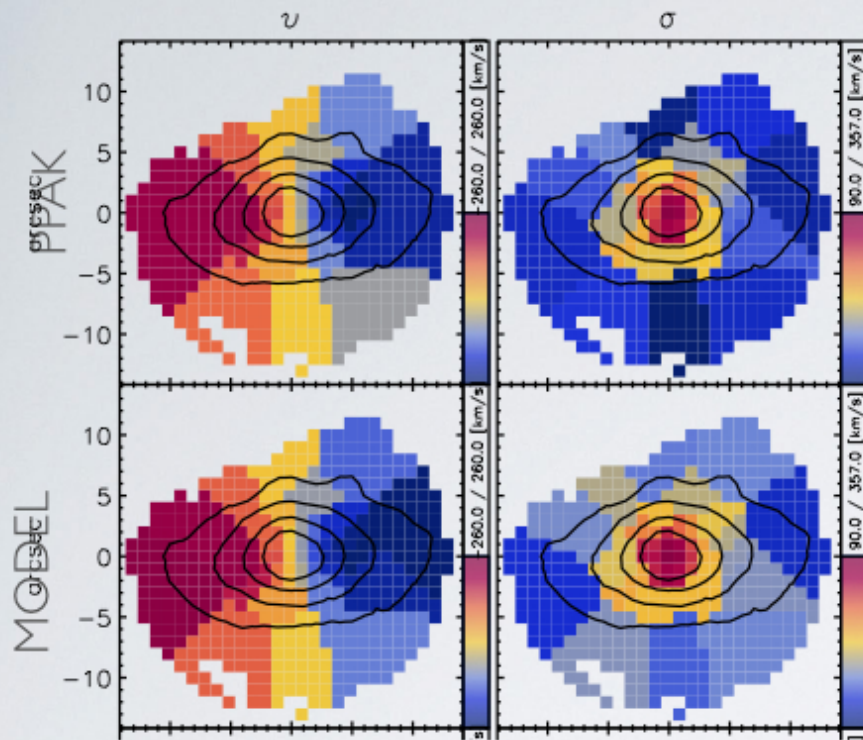


M32



**DRESSLER &
RICHSTONE (1988)**

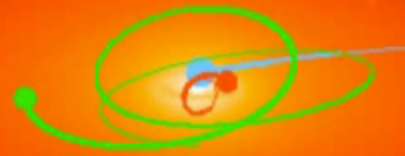




- PPAK IFU observations constrains the dynamical model and give a higher mass-to-light ratio and smaller black hole

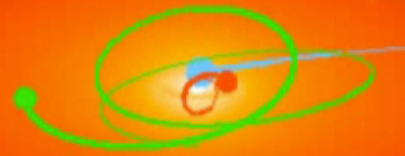


NGC 1277 IS AN OLD DISK GALAXY



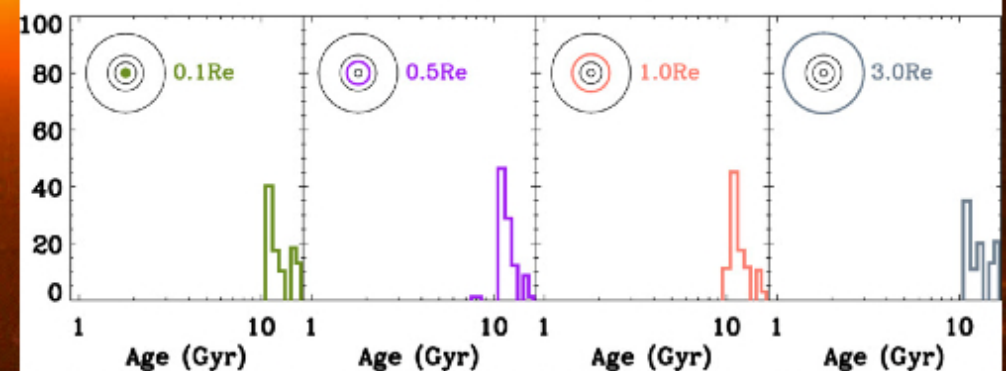
- No Classical Bulge, which implies no coevolution.
- Bottom heavy stellar populations (Emsellem 2013)
- stellar ages > 10 Gyr. (Trujillo+2014)
- Chandra X-ray luminosity of $1e40$ (Fabian et al. 2013), implies low accretion rate.
- That still leaves a lot of options... including Precipitation (Voit)

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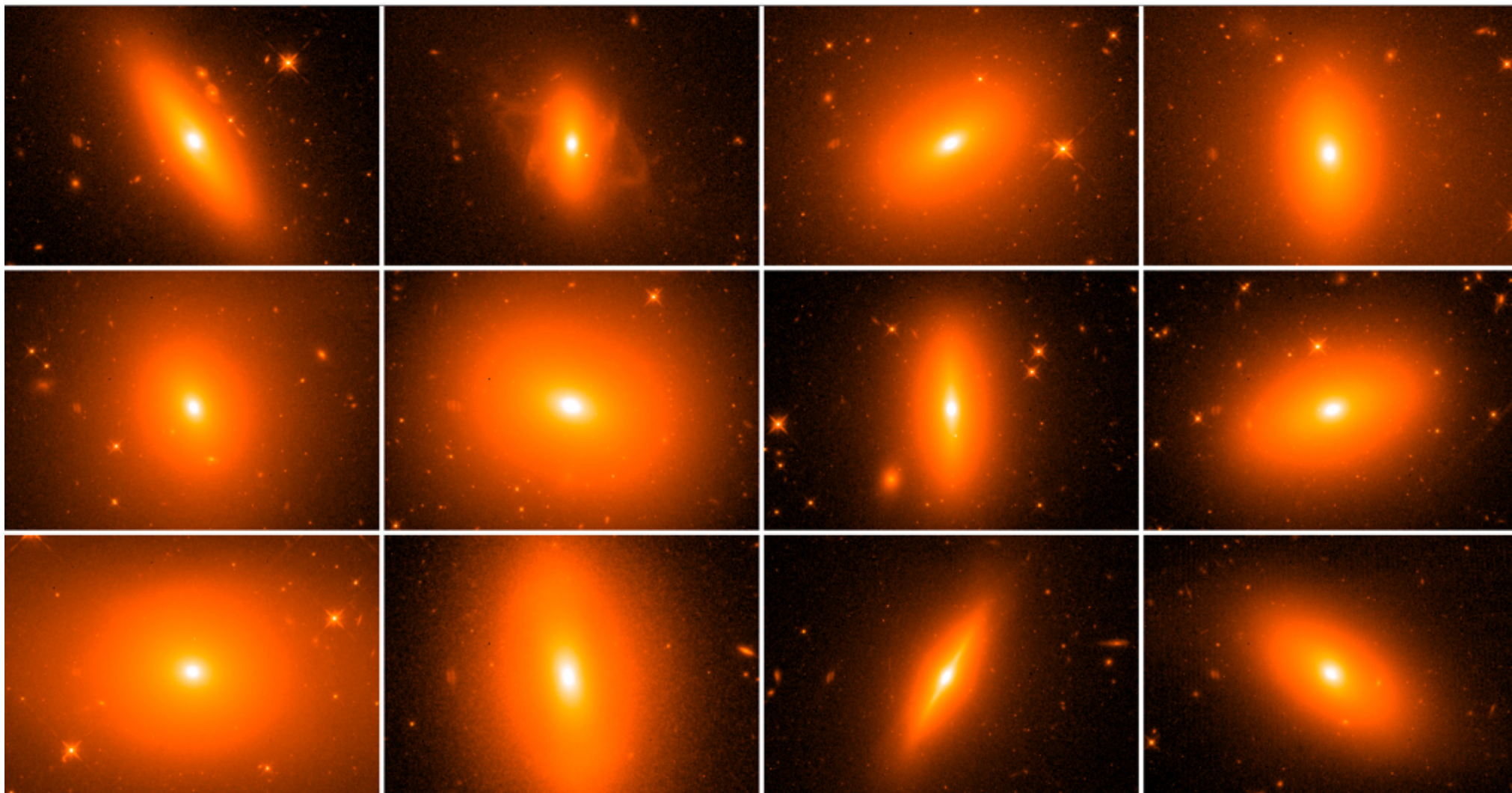
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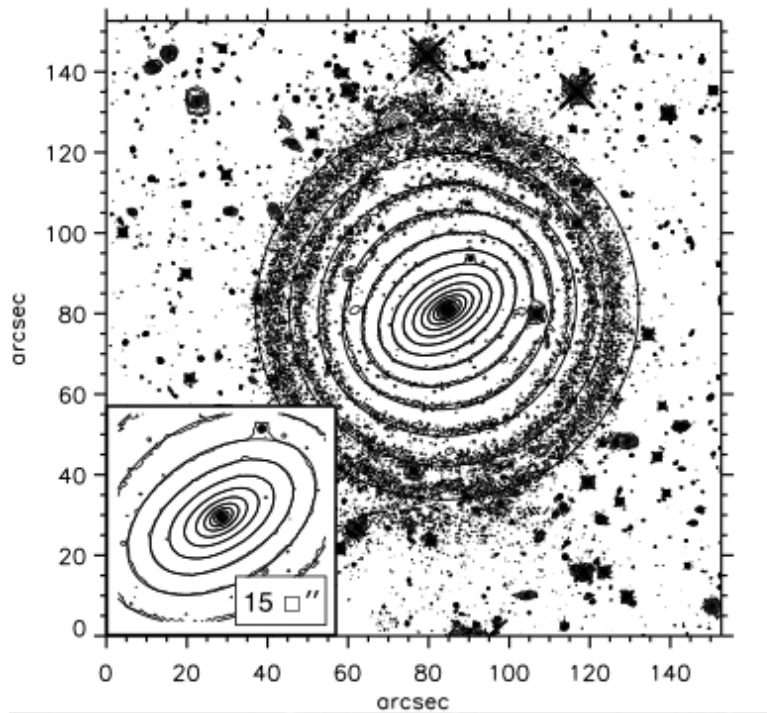
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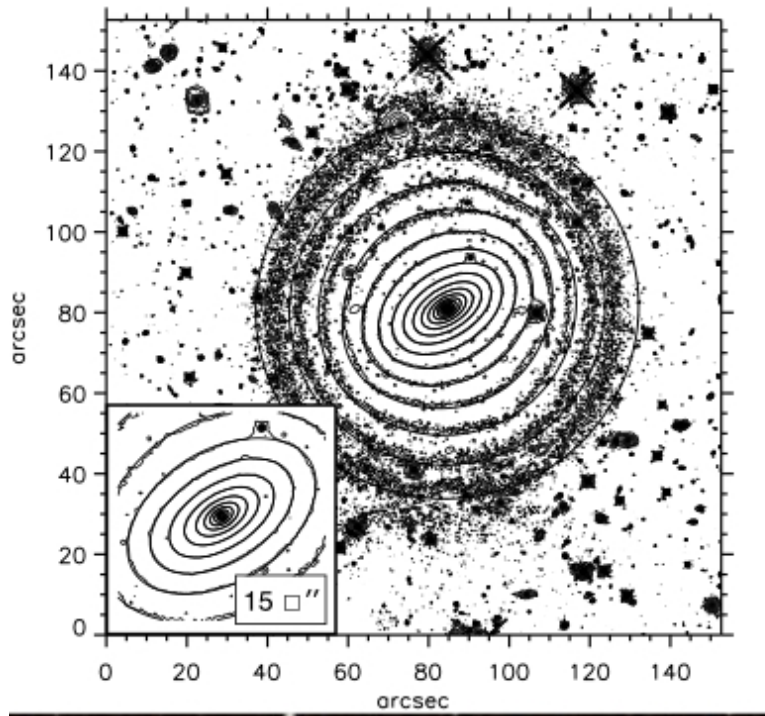
FOLLOW UP OF 17 COMPACTS WITH HST AND PPAK IFU



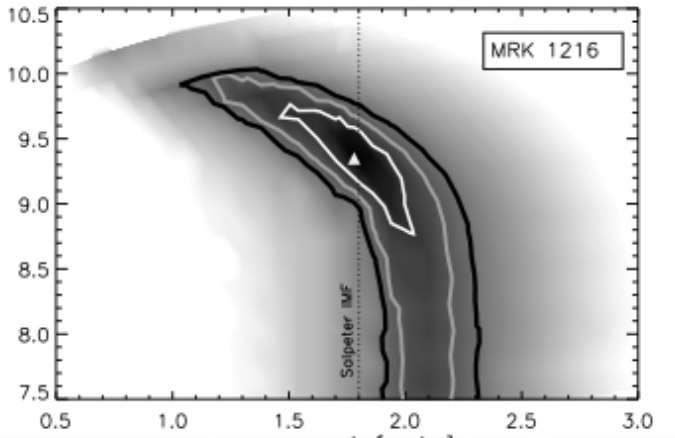


MRK1216



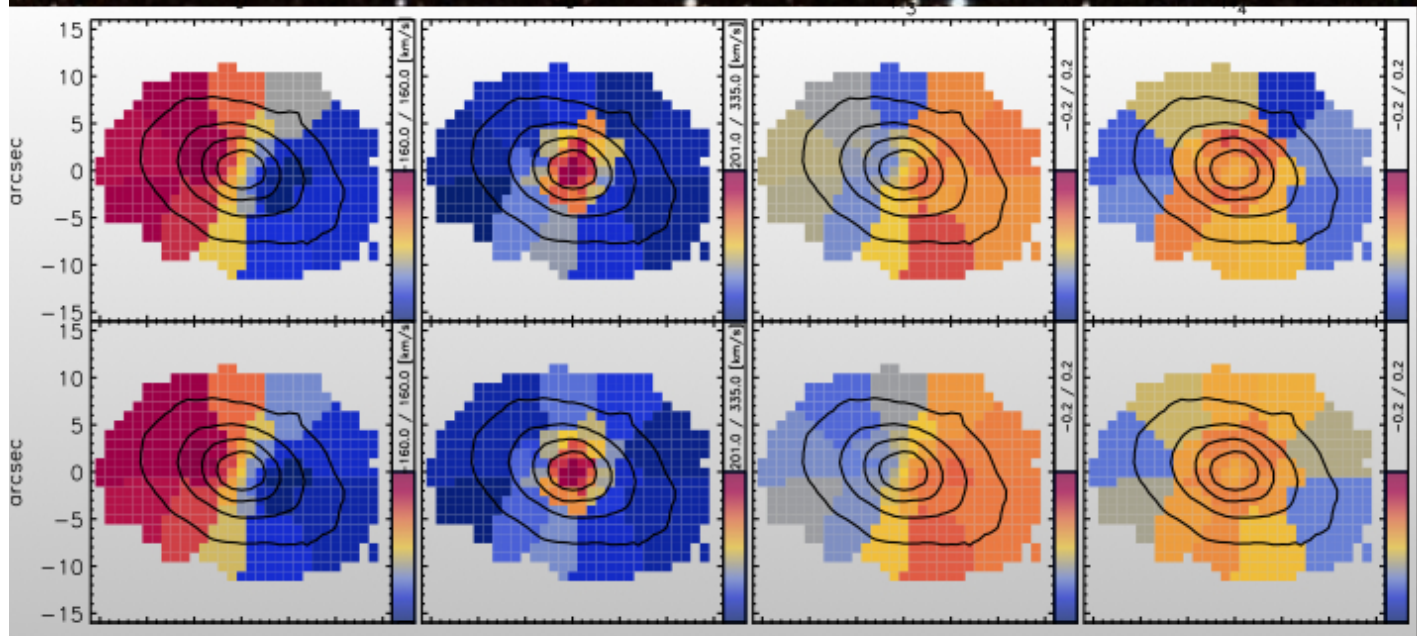


BH mass
log(Msun)

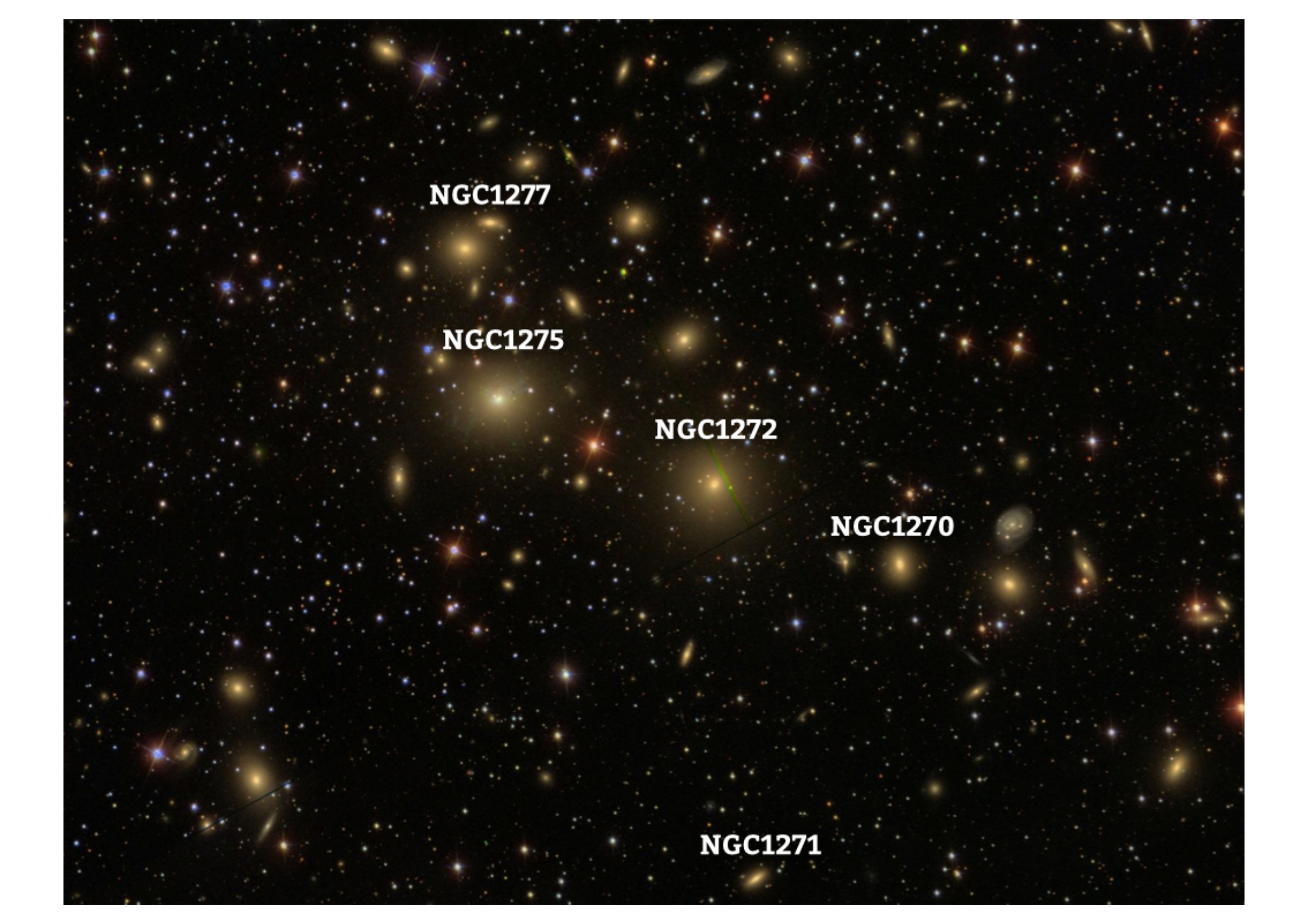


MRK1216

M/L_H



Yildirim+



NGC1277

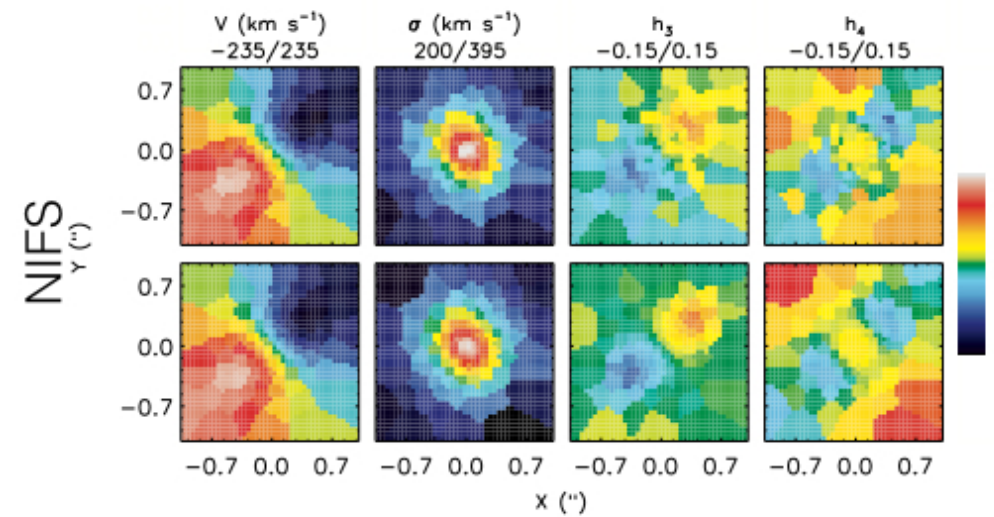
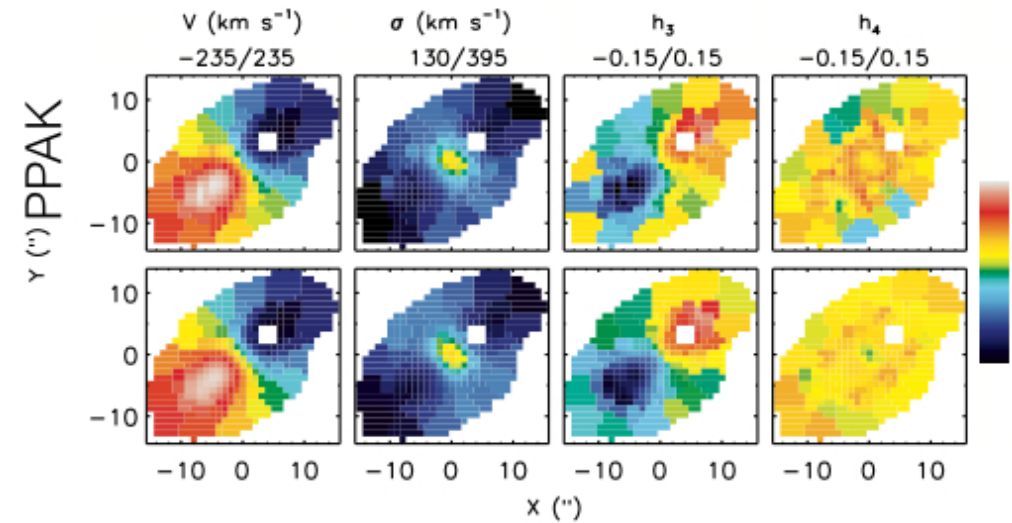
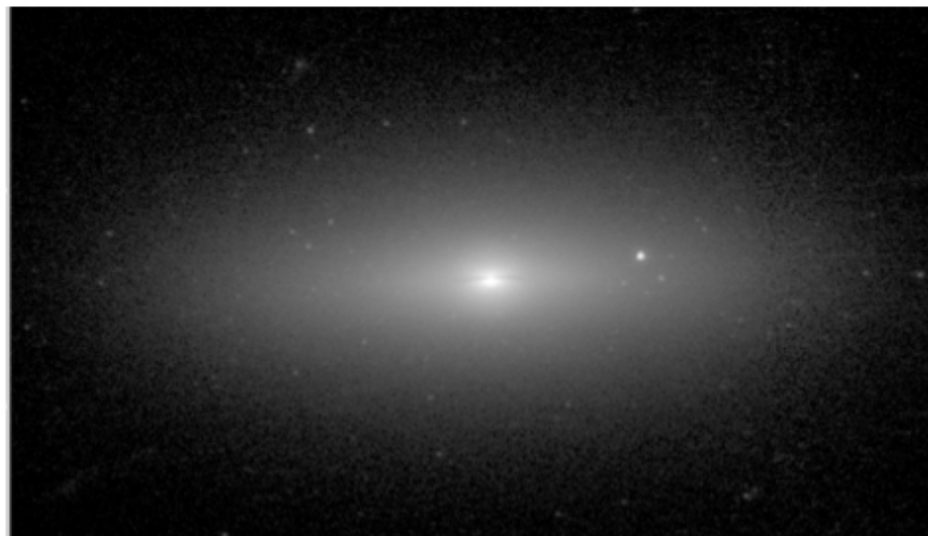
NGC1275

NGC1272

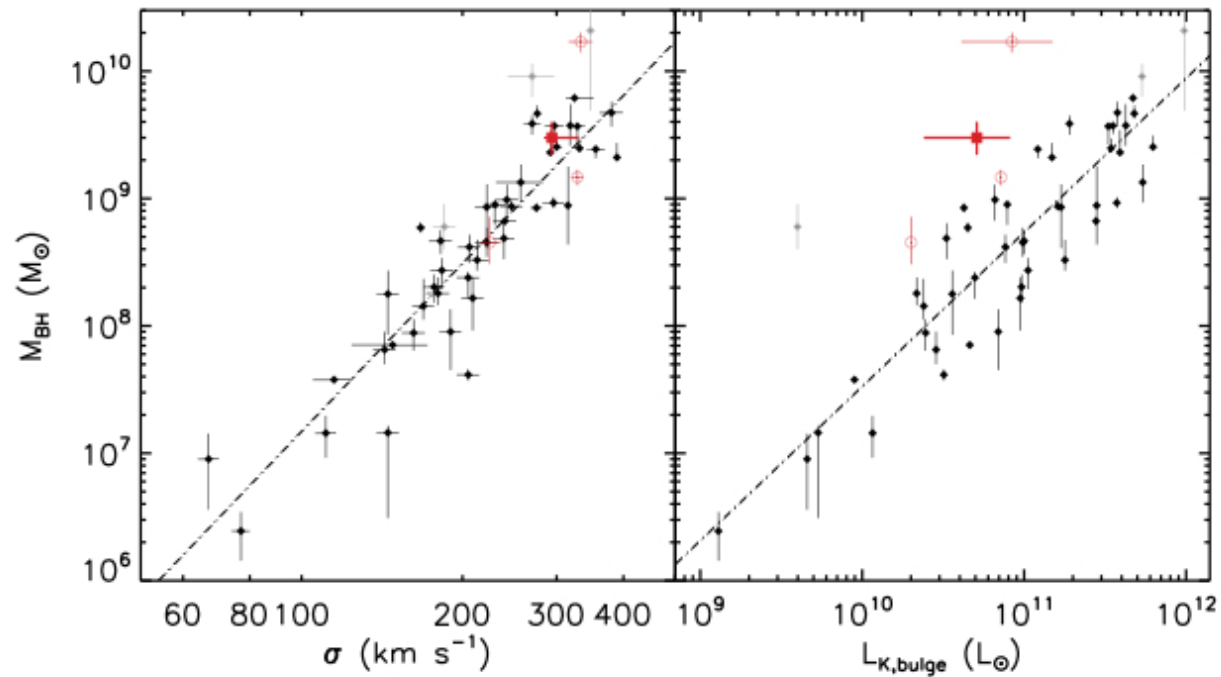
NGC1270

NGC1271

NGC1271 WITH PPAK AND NIFS

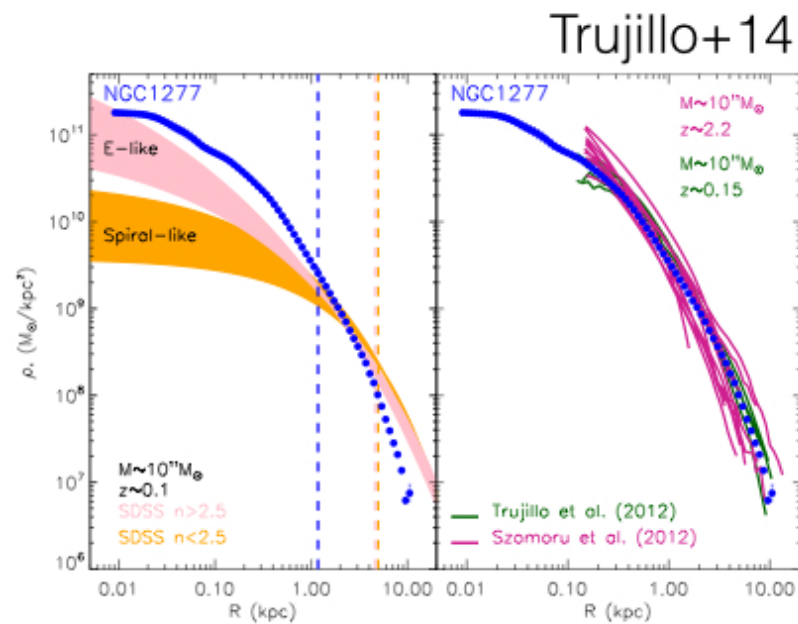


COMPACT GALAXIES

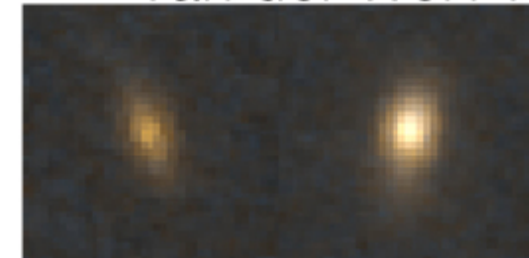


- Large lever on BH co-evolution

COMPACT GALAXIES



van der Wel+11



B/T = 0.25

B/T = 0.43

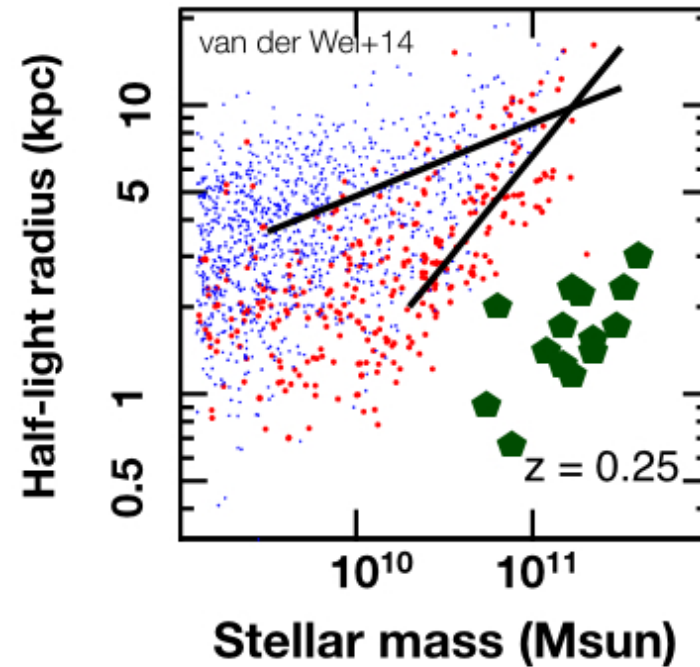


Rd = 1.7 kpc
Rb, eff = 0.33 kpc

Rd = 1.9 kpc
Rb, eff = 0.58 kpc

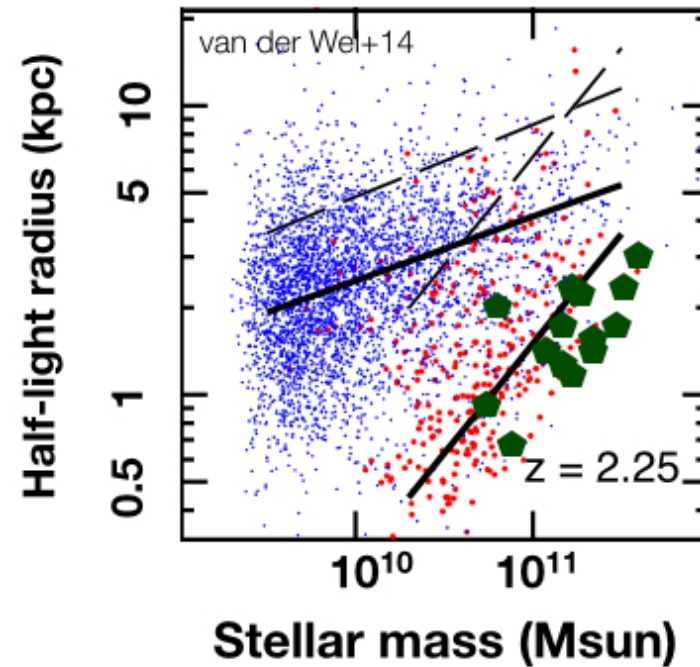
- Appear similar to $z \sim 2$ passive galaxies

COMPACT GALAXIES



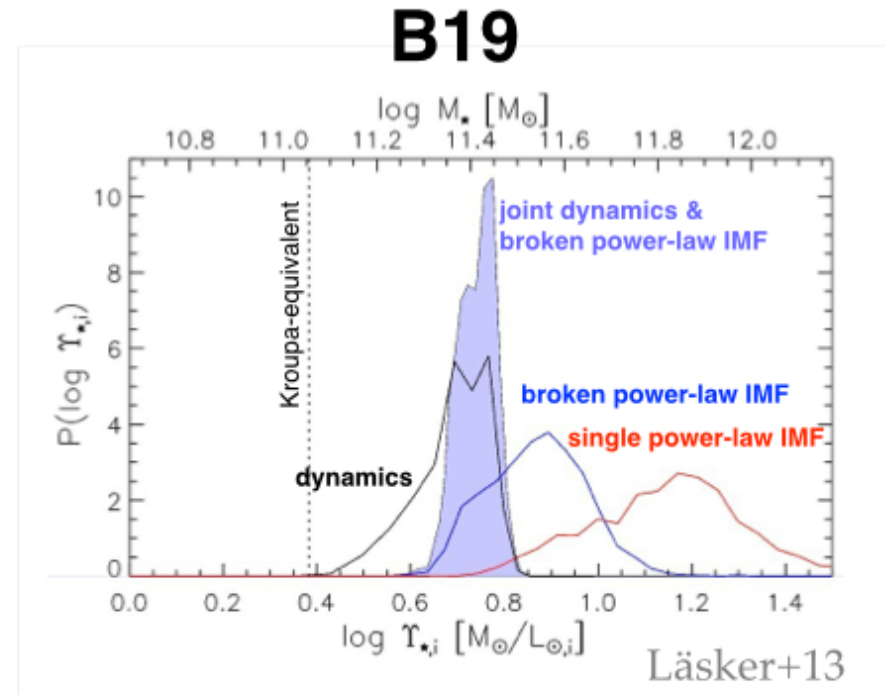
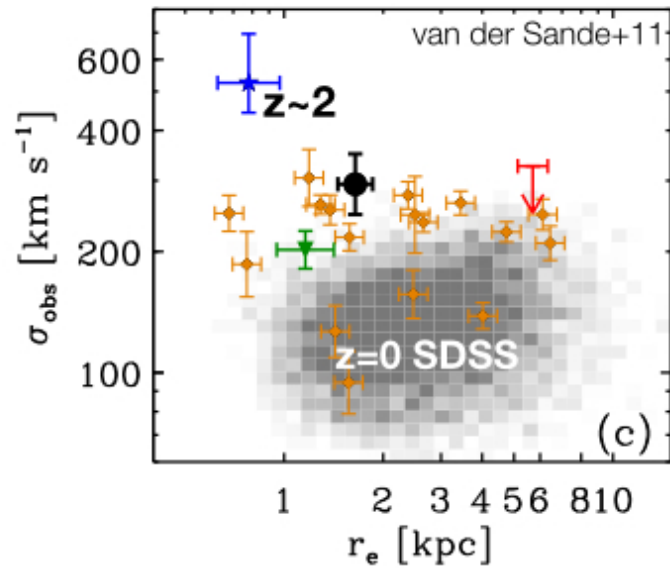
- Large mass-to-light ratios and hence bottom heavy IMF

COMPACT GALAXIES



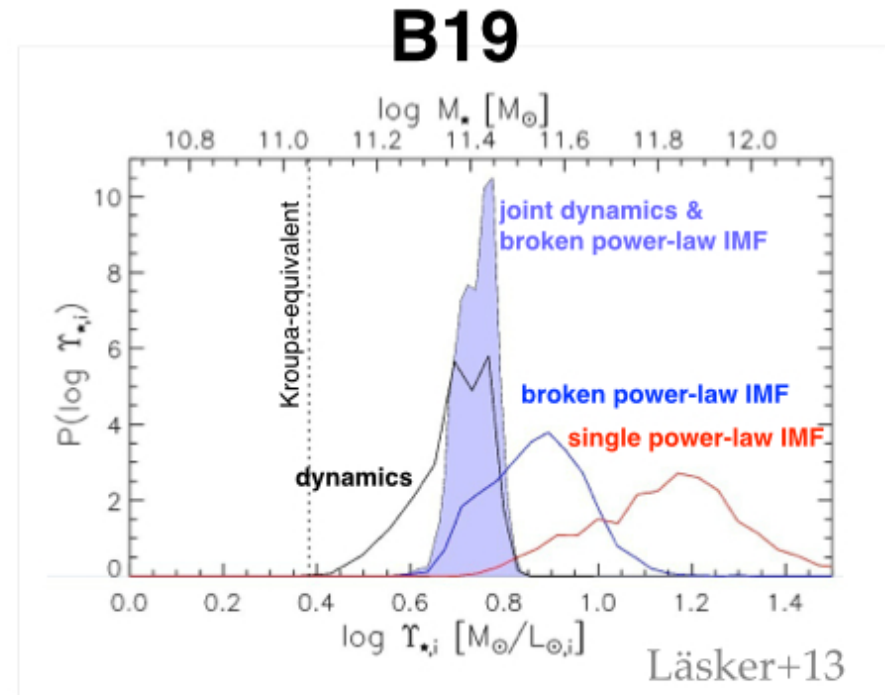
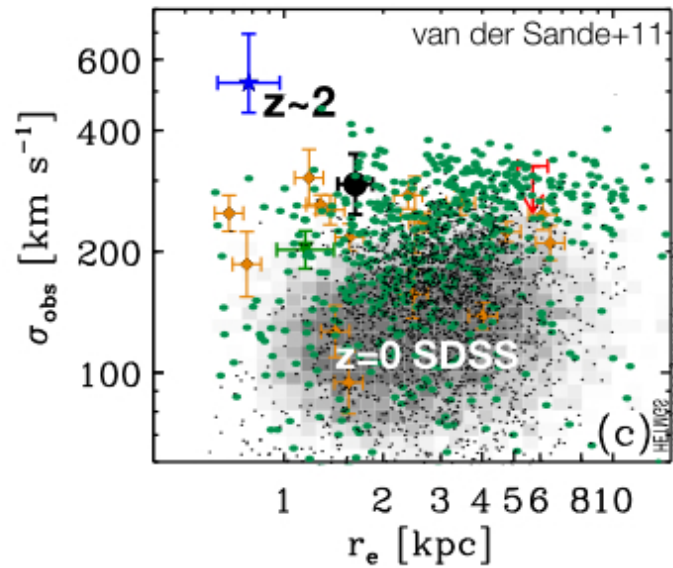
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COMPACT GALAXIES



- Exist in SDSS too!

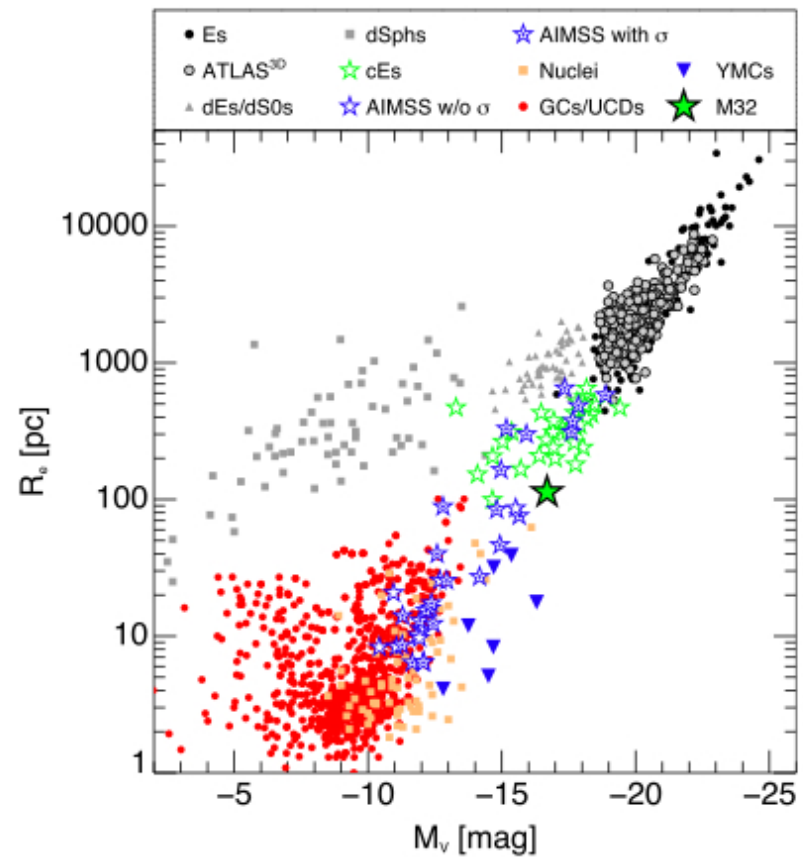
COMPACT GALAXIES



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Ultra Compact Dwarfs

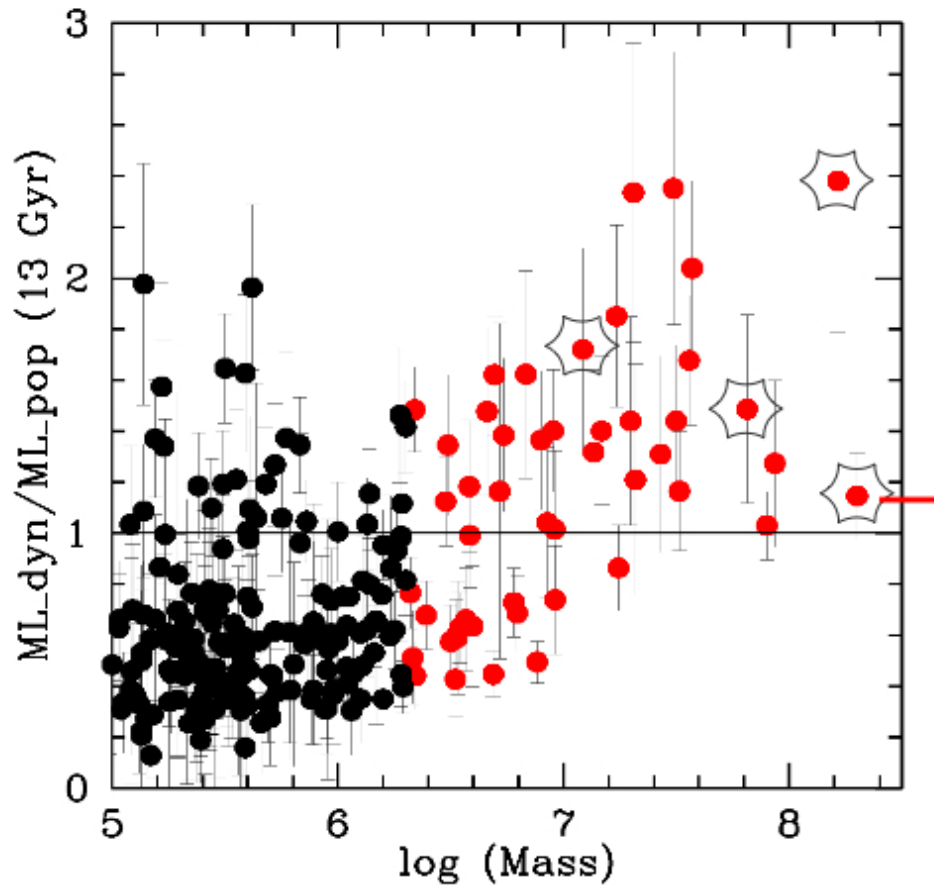
Big Globular Clusters or Stripped Galaxies?



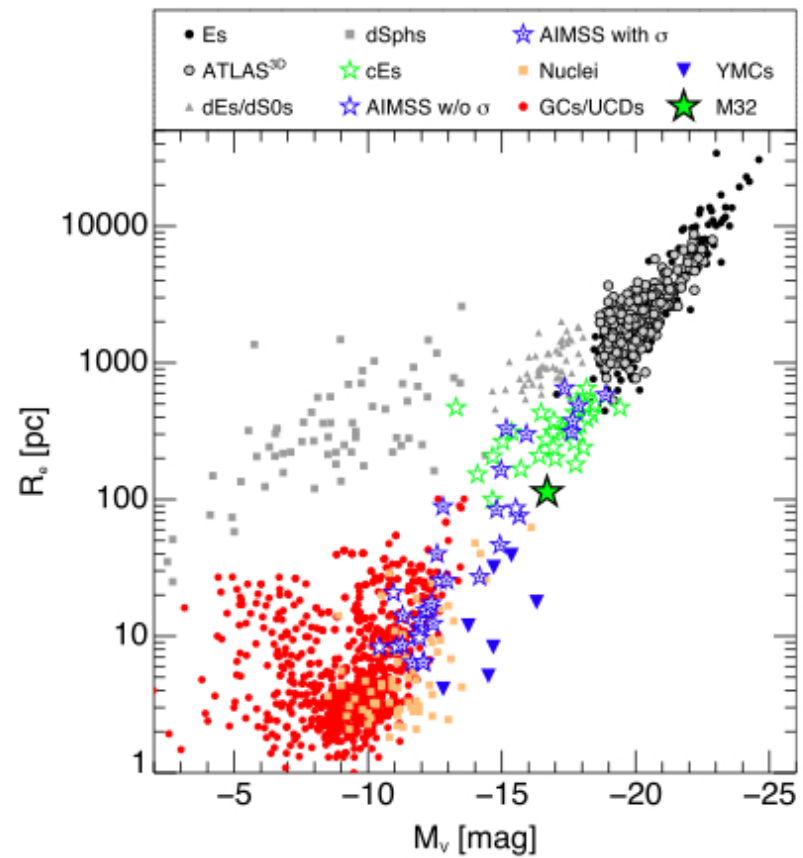
Norris+14

Ultra Compact Dwarfs

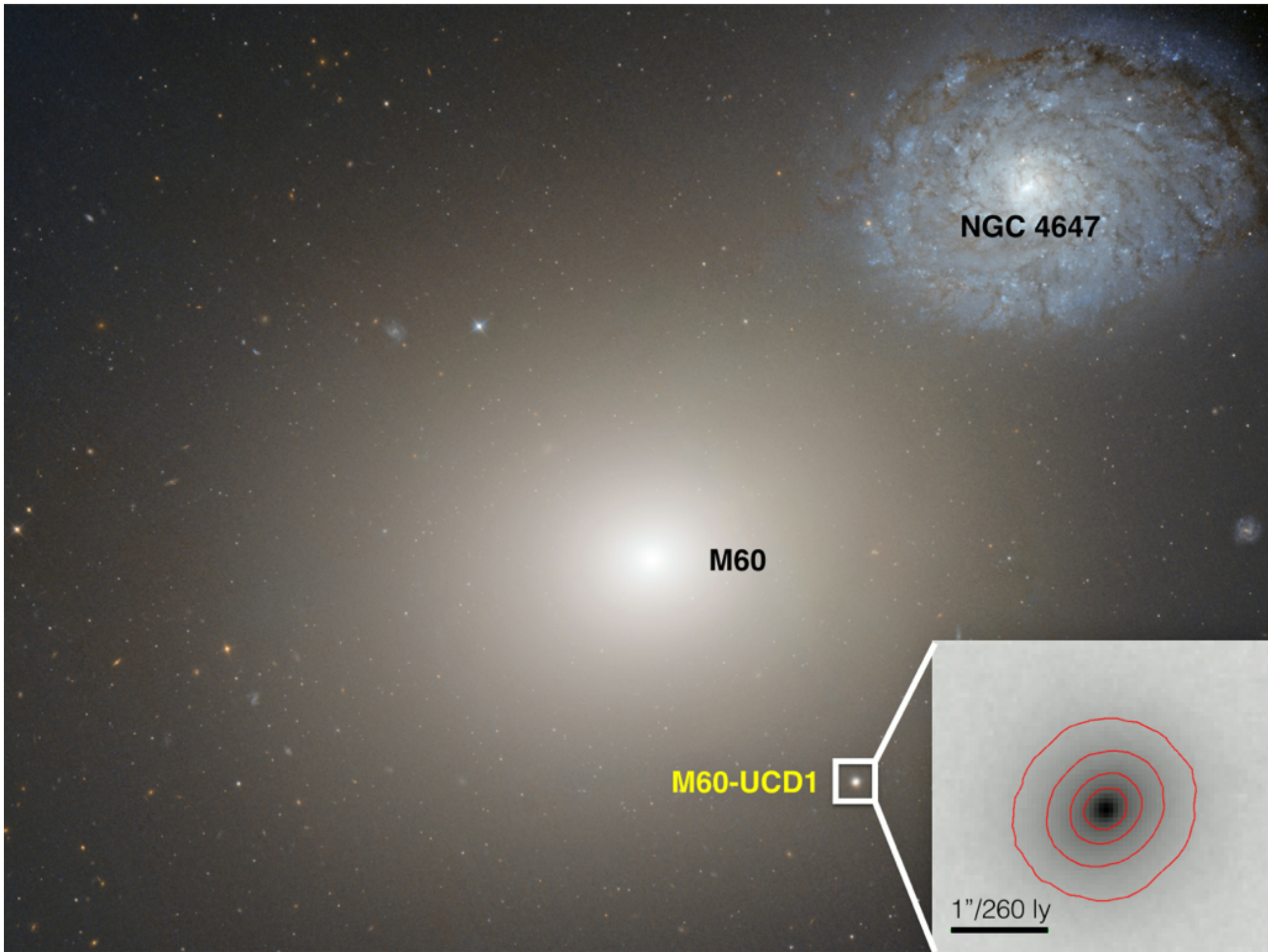
Big Globular Clusters or Stripped Galaxies?



Mieske+



Norris+14



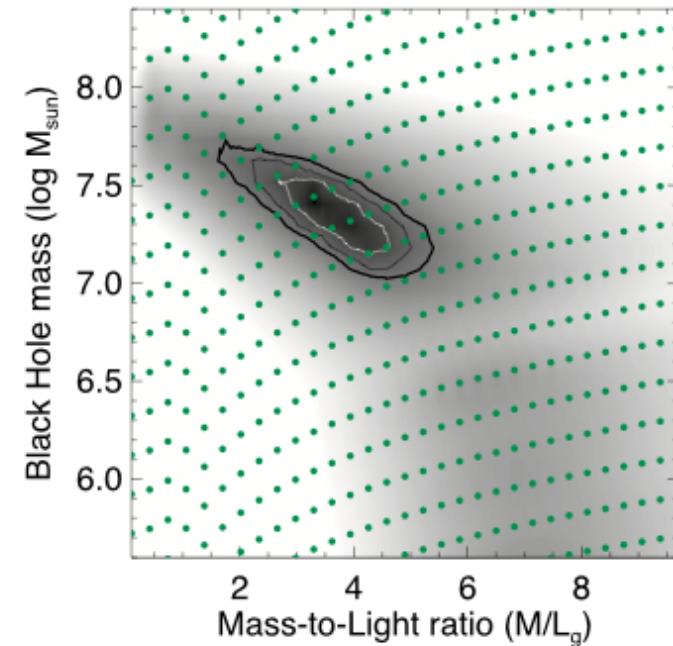
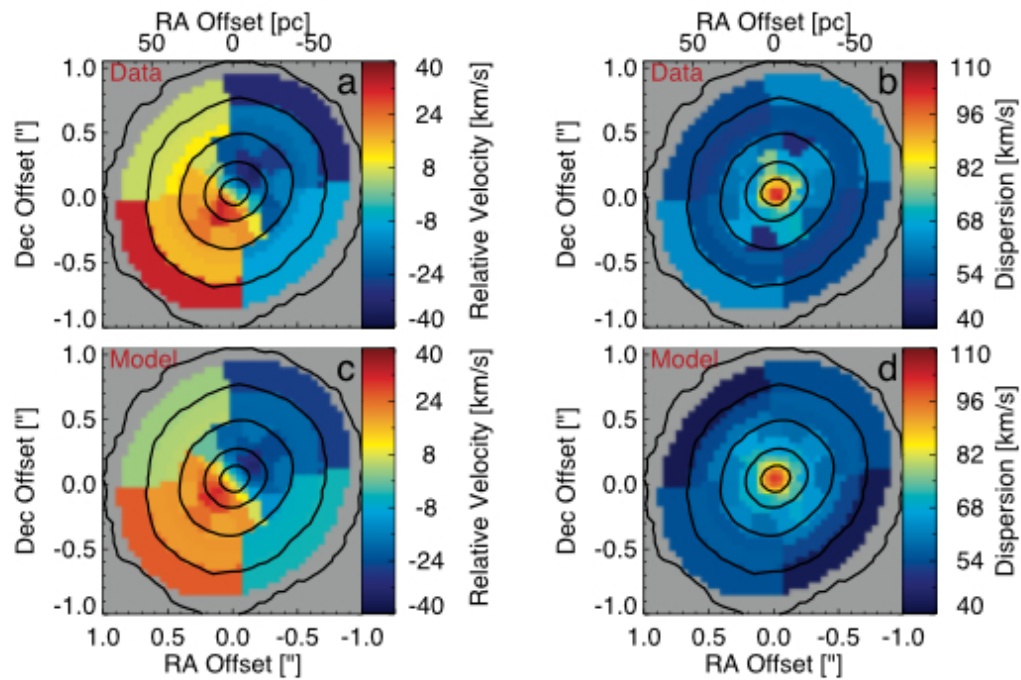
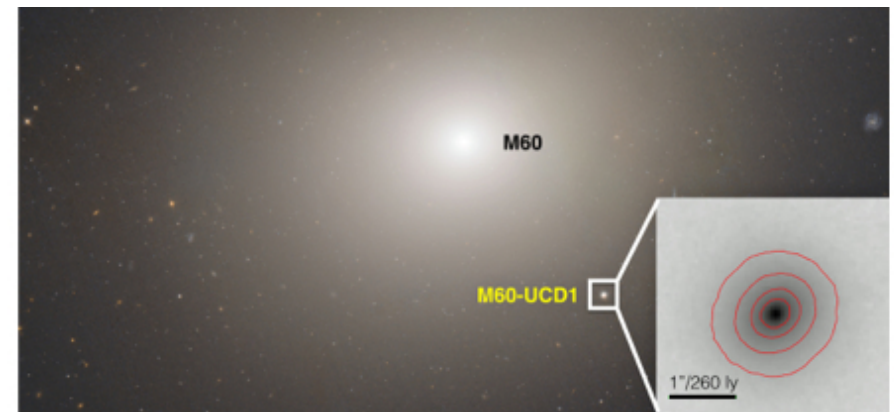
NGC 4647

M60

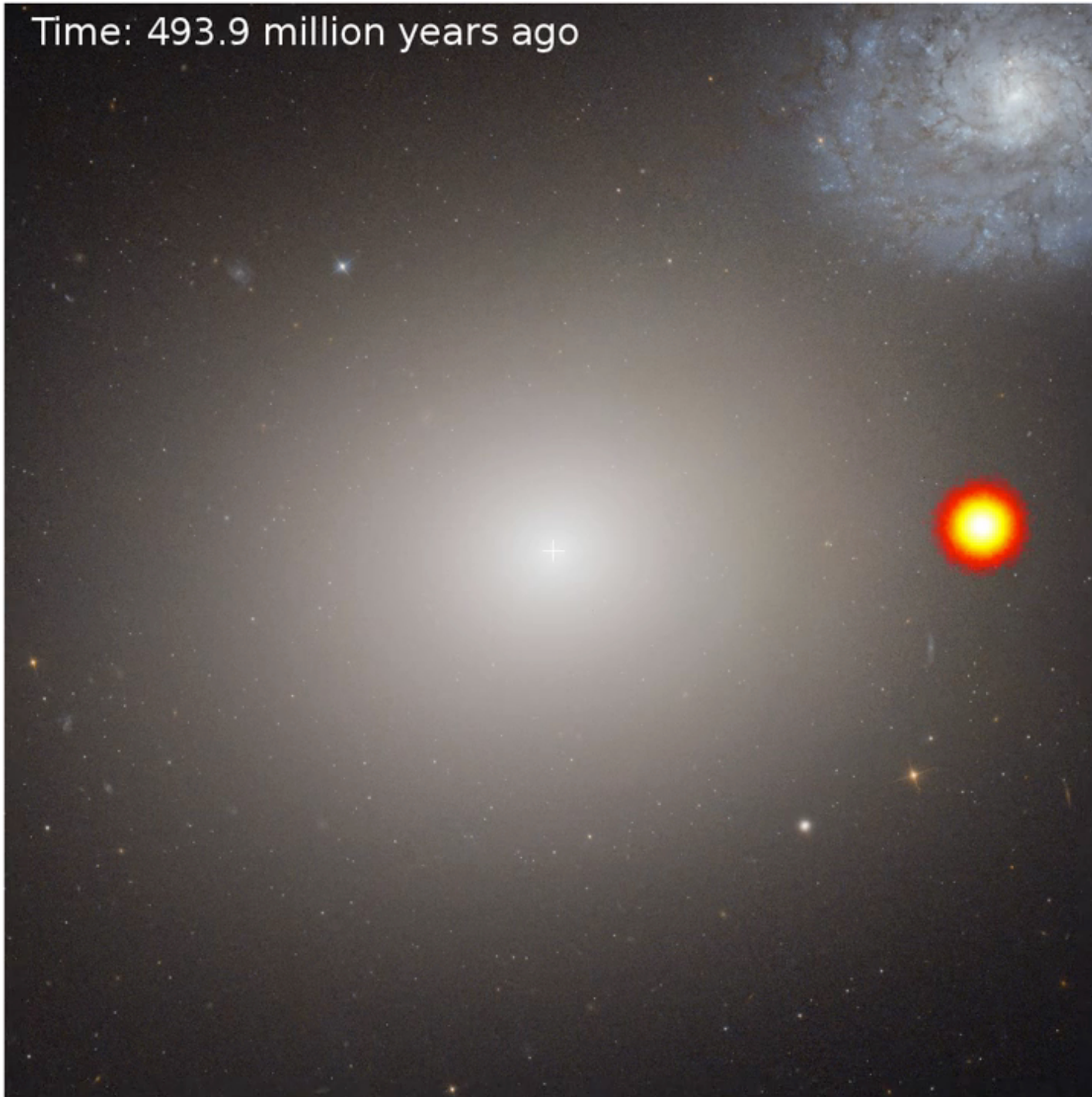
M60-UCD1

1"/260 ly

A super-massive black hole in an UCD



Time: 493.9 million years ago



CONCLUSIONS

- HET Massive Galaxy survey provides the necessary groundwork for future systematic black hole mass measurement campaigns.
- Compact Galaxies
 - Differentiate between different BH scaling relations
 - Appear very similar to $z \sim 2$ passive galaxies
 - Have large stellar mass-to-light ratio, which implies bottom heavy IMFs.